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19 20 21 22 23 24 25 26 27 28 29 30 31 32 33 34	The Eocene India-Eurasia collision is a first order tectonic event whose nature and chronology remains controversial. We test two end-member collision scenarios using coupled global plate motion-subduction models. The first, conventional model, invokes a continental collision soon after ~60 Ma between a maximum extent Greater India and an Andean-style Eurasian margin. The alternative scenario involves a collision between a minimum extent Greater India and a NeoTethyan back-arc at ~60 Ma that is subsequently subducted along southern Lhasa at an Andean-style margin, culminating with continent-continent contact at ~40 Ma. Our numerical models suggest the conventional scenario does not adequately reproduce mantle structure related to Tethyan convergence. The alternative scenario better reproduces the discrete slab volumes and their lateral and vertical distribution in the mantle, and is also supported by the distribution of ophiolites indicative of Tethyan intra-oceanic subduction, magmatic gaps along southern Lhasa and a two-stage slowdown of India. Our models show a strong component of southward mantle return flow for the Tethyan region, suggesting that the common assumption of near-vertical slab sinking is an oversimplification with significant consequences for interpretations of seismic tomography in the context of subduction reference frames.

## 35 **1. Introduction**

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37 The closure of the Tethyan ocean basins was responsible for the uplift of the vast Alpine-Himalayan 38 orogenic belt stretching from central Europe to Southeast Asia [Acharyya, 1998; Golonka, 2004; 39 Stampfli and Borel, 2002], with significant consequences for regional tectonics and global climate [Raymo and Ruddiman, 1992; Tapponnier et al., 1982]. The final stage of Tethyan 40 41 evolution involved the separation of India from Gondwana sometime after ~130 Ma [Robb et al., 42 2005], marked by magnetic anomaly M10 offshore Western Australia, to open the Indian Ocean at 43 the expense of the intermediary Meso- and Neo- Tethys [Veevers et al., 1991]. The Yarlung-44 Tsangpo Suture Zone (Figure 1) represents the main contact between the Indian and Eurasian continents where Indian epicontinental shallow-marine sediments juxtapose those of Eurasian 45 46 affinities [Najman and Garzanti, 2000; Searle et al., 1987; Yin and Harrison, 2000]. However, the 47 timing of the first mixing between Greater Indian and Eurasian sediments remains controversial, resulting in poorly constrained dynamics of the collision. We aim to use numerical models and 48 49 evidence from the subsurface to help constrain the nature and chronology of the India-Eurasia 50 collision. A better understanding of the convergence history has the potential to increase our 51 understanding of the timing of uplift in proximal orogenic belts, regional volcanic activity, 52 denudation histories and deformation.

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The established view of the collision is based on interpretations of magnetic lineations in the Indian Ocean seafloor that are used to derive India-Eurasia convergence histories [*Lee and Lawver*, 1995; *van Hinsbergen et al.*, 2011a]. Conventional interpretations link the reduction in India-Eurasia convergence rates and the first influx of India-derived sediments at the Eurasian margin to continent-continent collision sometime between ~65 and 50 Ma [*Garzanti*, 2008; *Lee and Lawver*, 1995; *Patriat and Achache*, 1984; *Replumaz and Tapponnier*, 2003; *Rowley*, 1996; *Searle et al.*, 1987]. We define "conventional models" as those that have long-lived Andean subduction along 61 southern Lhasa, with a terminal collision between the two continents occurring at the first significant slowdown of India-Eurasia convergence. An example of such models is one proposed by 62 63 Lee and Lawver [1995], which implies collision between a very large Greater India and southward 64 displaced Lhasa at ~55 Ma, forming the inspiration for one end-member scenario in our numerical 65 models. Alternative models now place the continental collision as late as 34 Ma [Aitchison et al., 66 2007]. Such models link the first slowdown in India-Eurasia convergence after 60 Ma to an initial 67 collision with a NeoTethyan intra-oceanic island-arc, followed by continent-continent collision 68 closer to ~40 Ma [Aitchison et al., 2007; Davis et al., 2002; Hafkenscheid et al., 2006; Klootwijk et 69 al., 1985; Van der Voo et al., 1999b]. We test two end-member plate kinematic scenarios of the 70 India-Eurasia collision to identify whether long-lived Andean style subduction at the Eurasian 71 margin implied by conventional models can account for the subducted slabs observed in seismic 72 tomography of the mantle, or whether a Meso- and Neo-Tethyan back-arc basin, proposed in 73 alternative models, is necessary to reproduce the present-day mantle structure. The kinematic 74 scenario that better reproduces the large latitudinal range of slab material at mid-mantle depths 75 interpreted from mantle tomography would therefore better represent the nature and chronology of 76 the India-Eurasia collision.

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## 78 1.1. Rotation models of India-Eurasia convergence

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The convergence history quoted by most authors to infer initial continental collision is usually based on a single plate motion model, with specific plate circuits and inherent interpretive biases in identifying magnetic lineations. Therefore, it is important to identify whether the choice of rotation models alone can account for the wide range in collision timing interpretation. Multiple plate kinematic models exist that document the relative motion between India and Eurasia, including those of *Lee and Lawver* [1995], *Müller et al.* [2008], *Molnar and Stock* [2009], and Model A of *van Hinsbergen et al.* [2011a]. When comparing the convergence rates of India to Eurasia across

87 published models (Appendix A and B), standardised to the geological time-scale of Cande and 88 Kent [1995], it is evident that the convergence-rate drops, and therefore inferences of initial 89 collisions, are entirely model-dependent and do not occur during the same time intervals with 90 different rotation models (Figure 2). The Lee and Lawver [1995] and Molnar and Stock [2009] 91 rotation models imply a convergence rate drop to less than 12 cm/yr at ~56 Ma, followed by other 92 stepwise decreases in convergence velocities at ~47 Ma and ~40 Ma, while the initial convergence 93 velocity decrease to values below 12 cm/yr occurs at ~55 Ma in Müller et al. [2008], and ~51 Ma 94 in van Hinsbergen et al. [2011a]. All of the rotation models suggest multiple stepwise decreases in 95 India-Eurasia convergence, and therefore it is difficult to identify a singular event related to India-96 Eurasia continental collision. The initial slowdown in seafloor spreading at ~52 Ma on the Central and Southeast Indian ridges, interpreted by Cande et al. [2010] using updated seafloor spreading 97 98 data and plate circuits, highlights that the choice of plate circuits and the identification of magnetic 99 lineations can significantly change plate velocity trends, meaning that conclusions based on 100 convergence rate trends alone should not be used to infer collision timing.

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## 102 1.2. Pre-collision margins

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104 Much of the collision timing controversy stems from the poor constraints on the pre-collision 105 margins of both Eurasia and Greater India [Aitchison et al., 2007; Hafkenscheid et al., 2006]. The 106 northward extent of Greater India in the Lee and Lawver [1995] model is derived from the 107 reconstructed position of India and Eurasia at ~55 Ma, at which time the gap between the two 108 continents is filled to ensure contact between Lhasa and India. However, this assumes that the ~55 Ma event, marked by a drop in the convergence rate in their rotation model, is a continent-109 110 continent collision, and that the active pre-collision Eurasian margin was an Andean-style margin 111 along southern Lhasa. In addition, microcontinent formation, whereby a passive continental margin 112 is rifted along a preferential landward detachment zone, would have successively removed

113 continental blocks from the Tethvan Gondwana margin [Müller et al., 2001] to reduce the size of 114 Greater India. Retro-deformation of the crumpled and under-thrust Indian crust below Tibet 115 suggests that the maximum northward extent of the passive Indian continental margin was no more 116 than 950 km of the present-day suture zone [Replumaz and Tapponnier, 2003], which agrees with the proposed smaller extent bound by the Wallaby-Zenith Fracture Zone at pre-breakup fit [Ali and 117 118 Aitchison, 2005] (Figure 4). A smaller Greater India is also compatible with the tomographically-119 derived pre-collision margin of Greater India of Replumaz et al. [2010] and the correlations of a 120 steeply-dipping unbroken slab in the upper mantle near the Yarlung-Tsangpo Suture Zone to the 121 subducted continental lithosphere of Greater India by Van der Voo et al. [1999b]. A range of 122 geometries have been proposed for this passive margin and are discussed at length in *Klootwijk and* 123 Conaghan [1979] and Ali and Aitchison [2005].

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125 In addition to the uncertain size of Greater India, the pre-collision margin of Eurasia is poorly understood. New paleomagnetic data, corrected for sediment compaction and inclination 126 127 shallowing [Tan et al., 2010], suggest that models such as Lee and Lawver [1995] place the Lhasa terrane, forming the active pre-collision margin, up to  $\sim 10^{\circ}$  too far south, indicating that early 128 129 contact at ~55 Ma between a maximum extent Greater India and southward-displaced Lhasa may be 130 problematic. Paleomagnetic studies which show an initial overlap between apparent polar wander 131 paths of India with respect to Asia suggest contact at  $46 \pm 8$  Ma [Dupont-Nivet et al., 2010] or as late as ~43 Ma [Tan et al., 2010], indicating that collision timing derived from paleomagnetic data 132 133 has large inherent uncertainties and cannot be used alone to define the initial timing of continent-134 continent collision. Reconstructing the geometry of Greater India and Lhasa as proposed by Lee and Lawver [1995] shows that initial contact occurs between 54 and 49 Ma across the five different 135 136 rotation models (Table 1 highlighted column and Figure 3), which highlights the model-dependence of interpretations based on rotation models to infer collision timing (Figure 2). When reconstructing 137 138 the maximum extent of Greater India from Lee and Lawver [1995] and a range of proposed paleomargins of Eurasia, the effect on the collision timing is much more pronounced than varying the rotation model alone, resulting in a collision window of 15 Myr (Figure 3, Table 1). Thus the choice of pre-collision margins of both India and Eurasia is the dominant variable determining the timing of continental contact in plate reconstructions.

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144 Independent of the rotation models and inferences of pre-collision margins, stratigraphic studies 145 focused on the horizons that mark the first input of Tethyan sediments overlying sediments of 146 Eurasian affinities to date the onset of continent-continent collision older than ~50 Ma [Clift et al., 147 2002]. However, a re-interpretation of this horizon by *Henderson et al.* [2011] shows only a break 148 in Eurasian sedimentation, with India-derived sediments occurring higher in the sedimentary column, indicating that continental collision may have occurred later than the ~50 Ma age proposed 149 by Clift et al. [2002]. Recent studies of sedimentary sequences and ophiolite belts near the main 150 151 suture zone suggest a well-established intra-oceanic subduction system in the NeoTethys whose closure preceded the main continent-continent collision [Aitchison et al., 2007; Aitchison et al., 152 2000; Davis et al., 2002; McDermid et al., 2002; Ziabrev et al., 2004]. Ophiolites of remnant back-153 154 arcs in the western NeoTethys have been studied extensively, including the Semail suprasubduction zone ophiolite that was emplaced after 90 Ma, and it has been suggested that a similar 155 156 scenario existed in the central NeoTethys [Hafkenscheid et al., 2006; Pearce et al., 1981; Shervais, 157 2001; Stampfli and Borel, 2002]. Cessation of arc-magmatism along the Kohistan-Ladakh arc at ~61 Ma can be interpreted as the onset of ophiolite emplacement onto Greater India [Khan et al., 158 159 2009], while further east the Zedong Ophiolite was likely to have been emplaced at ~57 Ma [Ali 160 and Aitchison, 2008]. The precise dating of ophiolite emplacement onto Greater India is 161 controversial because it relies on dating the first appearance of serpentinite-rich ophiolitic sediment 162 or paleobiological methods. It is also difficult to constrain the southward extent of the proposed 163 intra-oceanic magmatic arc using paleomagnetic data, with inherent shortcomings and typical

164 latitudinal error margins of  $\pm 5^{\circ}$  for the region [*Sun et al.*, 2010] that are as large as the proposed 165 back-arc basin itself.

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#### 167 1.3. Seismic tomography and numerical models of the collision

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169 In light of the complex geology, some authors have looked to the sub-surface for independent clues 170 of Tethyan subduction histories. The distribution of positive seismic velocity anomalies identified 171 in P-wave seismic mantle tomography north of the equator and beneath the collision zone have been 172 linked to long-lived Tethyan subduction evolution [Hafkenscheid et al., 2006; Van der Voo et al., 1999b]. Van der Voo et al. [1999b] link the latitudinal distribution of discrete slabs in the lower 173 174 mantle to paleo-subduction zone locations, and suggest that a subducted NeoTethyan back-arc basin 175 better accounts for the lateral and vertical distribution of slab material. The existence of an intra-176 oceanic subduction system would have increased the southward extent of the Eurasian active margin. Consequently, the subducted NeoTethyan slab should be found south of the present-day 177 178 Yarlung-Tsangpo Suture Zone [Hafkenscheid et al., 2006] assuming little lateral migration of slabs 179 following subduction, following Van der Meer et al. [2010], and the temporal persistence of slab-180 derived thermal anomalies in the mantle [Jarvis and Lowman, 2005; Van der Voo et al., 1999a].

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182 The quantitative analysis by Hafkenscheid et al. [2006] compared expected Tethyan slab volumes 183 from plate reconstructions with positive seismic velocity anomalies of 0.2% and higher from the P-184 wave tomography model of Bijwaard et al. [1998], indicating that long-lived NeoTethyan 185 subduction along an Andean-style margin does not adequately account for the present-day mantle 186 structure. Their scenario of intra-oceanic subduction along a large back-arc basin better reproduces 187 the volume and distribution of positive seismic velocity anomalies in the mantle. In their model 188 India collides with an intra-oceanic island arc at ~65 Ma and Andean-style subduction is 189 subsequently initiated along continental Eurasia to consume the back-arc basin by ~48 Ma *[Hafkenscheid et al.*, 2006]. *Jarvis and Lowman* [2005] tested the viability of such subduction
scenarios using 2D numerical subduction models of the India-Eurasia collision, identifying the need
to independently reproduce and quantify the observed Tethyan mantle seismic velocity anomalies
using numerical subduction models.

#### 197 **2. Method**

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## 199 2.1. Kinematics of the India-Eurasia collision

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201 Our global plate kinematics are based on the tectonic model of Müller et al. [2008] and Seton et 202 al. [submitted] and were used to generate the seafloor age-grids and plate velocities in 1 Myr 203 intervals (Figure 5). Intersecting plate boundaries through time were implemented in 204 GPlates (www.gplates.org) to define continuously closing plate polygons that cover the entire 205 globe, modified from *Gurnis et al.* [2012]. We use a moving hotspot absolute reference frame from 206 100 Ma [O'Neill et al., 2005] and true-polar wander-corrected motions of Africa form 100 to 207 200 Ma [Steinberger and Torsvik, 2008] in order to isolate the plate-mantle system. The fixed 208 hotspot absolute reference frame of Müller et al. [1993] was also considered, but the limited 209 temporal span of ~130 Myr made it unsuitable for our global models. In addition, the Van der Meer 210 et al. [2010] subduction reference frame was not used as it would result in circular reasoning while 211 imposing the assumption of vertical sinking of all slabs with a constant sinking rate. When 212 extending the finite rotations of the fixed hotspot reference frame (FHS) to 200 Ma, the reconstructed paleo-margin of Eurasia is offset ~10° northward at 200 and 150 Ma when compared 213 214 to the moving hotspot and true-polar wander corrected frame of reference in our global models 215 (Figure 6). The subduction reference frame results in a longitudinal offset of the paleo-margin at 216 these times due to the inherent longitudinal correction in this reference frame. At 100 and 50 Ma all 217 reference frames produce a similar paleo-reconstruction of the Tethyan and Pacific active margins, 218 with differences that we assume are not significant enough to be detectable in seismic tomography.

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A scenario broadly similar to that of *Lee and Lawver* [1995] with a large Greater India and longlived Andean-style subduction in the NeoTethys at the continental margin was implemented as the conventional model. The southward extent of undeformed Eurasia was based on the geometry 223 proposed by *Replumaz et al.* [2004] and the pre-collision southward extent of Lhasa based on *Lee* and Lawver [1995] (Figure 5, left column). The size of Greater India and the NeoTethyan intra-224 225 oceanic subduction of the alternative kinematic scenario are largely based on the preferred subduction model of Hafkenscheid et al. [2006], and a scenario slightly modified from that 226 227 proposed by Ziabrev et al. [2004], Ali and Aitchison [2008] and Aitchison et al. [2011] (Figure 5, 228 right column). The rotations of India and Eurasia were not modified, and only the pre-collision 229 margins were varied. The smaller Greater India is based on the interpretations of Replumaz et 230 al. [2010] and Ali and Aitchison [2005] (Figure 4).

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## 232 2.2. Back-arc extent from age-coded slabs in seismic tomography

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234 The extent of the proposed NeoTethyan intra-oceanic subduction zone in the alternative scenario 235 was constructed by determining the approximate location of paleo-subduction using age-coded depth slices of subducted material in P- and S- wave seismic tomography models (Figure 7, 236 237 Appendix C). We interpret positive seismic velocity anomalies in the mantle that correlate across 238 both P- and S- wave tomography models to be slab remnants representing thermally-perturbed 239 mantle [Becker and Boschi, 2002]. Following Hafkenscheid et al. [2006], we assumed an average 240 sinking rate of 3 cm/yr in the upper mantle for the Tethyan subduction system. For the lower 241 mantle, we assumed a sinking velocity of 1.2 cm/yr from the global tomography interpretations of 242 Van der Meer et al. [2010]. We assume purely vertical sinking of slabs in the mantle for this 243 workflow, and we use our geodynamic models to determine the plausibility of such assumptions. In 244 the 1604 km depth slice from the Li et al. [2008] P-wave model (MIT-P), slab material which corresponds to  $\sim 100$  Ma with the assumed sinking velocities, shows a large mismatch of up to  $\sim 20^{\circ}$ 245 246 latitude between the *Replumaz et al.* [2004] Eurasian continental margin and the slab material in the 247 depth slice (Figure 7). The 1658 km depth slice in the Grand [2002] S-wave model (GRAND-S), 248 corresponding to ~105 Ma, shows a similar mismatch. We used the southern boundary of the 249 positive seismic velocity anomaly as the maximum southward extent of the NeoTethvan back-arc 250 basin. We then created rotations and isochrons to simulate the opening of a back-arc basin and to 251 generate seafloor age-grids consistent with the alternative scenario. As the Lhasa terrane accreted 252 onto Eurasia at the Jurassic-Cretaceous boundary [Golonka, 2004], we assumed that NeoTethyan back-arc basin opening occurred between 150 and 120 Ma with a full spreading rate of about 253 254 5.5 cm/yr and a maximum north-south extent of 1600 km. The back-arc closes only following 255 collision with the Greater Indian margin at about 60 Ma, with back-arc subduction and continental 256 suturing occurring by ~40 Ma. Although a mechanism for initiating intra-oceanic subduction in the 257 Tethys is enigmatic, we assume that the intra-oceanic subduction system could develop from 258 possible MesoTethyan slab rollback, back-arc rifting and spreading following the welding of the 259 Cimmerian terranes onto southern Eurasia in the late Jurassic.

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#### 261 2.3. Global subduction model setup

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263 As plate motions are the surface manifestation of mantle convection [Bercovici et al., 2000], forward numerical models that couple plate velocities with subduction allow us to predict present-264 day mantle structure that can be validated using seismic tomography [Ricard et al., 1993]. We used 265 266 a modified version of the finite element code CitcomS [Tan et al., 2006; Zhong et al., 2000] to solve for thermal convection within an incompressible viscous mantle with plate velocities applied as 267 268 kinematic boundary conditions on the upper surface. The subduction models used a mesh with 269 lateral grid spacing of 50 and 23 km on the surface and at the core-mantle boundary (CMB) 270 respectively (Table 2). To balance the reduced resolution closer to the surface, we refined the 271 vertical resolution of the mesh in the upper mantle and used a global domain to avoid edge effects 272 that would normally occur in regional models. A half-space cooling model for the oceanic lithosphere was progressively assimilated into the numerical models, following Matthews et 273 al. [2011], using the seafloor age-grid from our plate reconstructions. The continental lithosphere is 274 275 divided into three categories: Archean, Proterozoic and Phanerozoic [Artemieva, 2006] for which

we impose a thermal lithosphere thickness of 215, 165 and 100 km respectively [*Artemieva*, 2009]. The thermal structure of slabs is imposed at each timestep with assimilation stencils that extend to a distance of 350 km around a subduction zone, and a 200 km curvature radius of subducting slab material (Gurnis, in prep.). We use the parameters in Table 3 to define our Rayleigh number for mantle depths,  $Ra = -4 \times 10^7$ , and we apply temperature-dependent viscosity following

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$$\eta = \eta_0 \times \exp\left(\frac{E_\eta}{T^* + T_n} - \frac{E_\eta}{0.5 + T_n}\right)$$

where  $\eta_0$  is the reference viscosity,  $E_\eta$  is the activation energy,  $T^* = \min(\max(T, T_{min}), T_{max})$ , with *T* as the temperature,  $T_{min}$  and  $T_{max}$  the minimum and maximum temperatures of the mantle, respectively, and  $T_\eta$  the activation temperature (Figure 8, Table 3). We ignore the effects of internal heating from radioactive decay, but allow the bottom thermal boundary layer to develop dynamically through basal heating with a CMB non-dimensional temperature of 1. A viscosity contrast of 100 between the upper and lower mantle was implemented in our model runs that is consistent with the findings of *Alpert et al.* [2010] and *Jarvis and Lowman* [2005; 2007].

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# 290 **2.4 Comparison to seismic tomography**

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A number of mantle seismic tomographic models were chosen to compare with the predictions of 292 293 the numerical models. As each seismic tomography model has inherent assumptions, variable coverage, vertical and lateral resolutions, and damping parameters [Grand, 2002; Romanowicz, 294 295 2008], we consulted a number of models. Although P-wave models tend to have higher resolutions and better-resolved subduction zones, coverage for oceanic regions is limited [Grand, 2002]. 296 297 Longer wavelength features in the mantle are better imaged with S-wave models, which also tend to 298 have better coverage in the southern hemisphere and in the Pacific [Romanowicz, 2008]. We do not 299 attempt to analyse the mantle structure below 2500 km depth which likely has significant chemical 300 heterogeneities [Masters et al., 2000], and as both kinematic scenarios invoke subduction in the 301 northern hemisphere, we avoided analysing southern hemisphere mantle structure due to lower

- 302 resolution seismic tomographic models. A description of all tomographic models used can be found
- 303 in Appendix D.

- 308 **3. Results**
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#### 310 3.1. Vertical subduction zone evolution

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312 A vertical slice along a great circle segment between 10°S, 55°E and 45°N, 90°E was chosen for 313 comparison between the conventional and alternative scenario as it was the most representative of 314 the NeoTethyan convergence and proposed back-arc evolution. Plate kinematics between 200 and 315 150 Ma are common to both scenarios (Figure 9a), with continuous Andean-style subduction along the Eurasian margin. The subduction zone retreats almost 10° to the south between 200 and 150 Ma 316 317 due to the clockwise rotation of Eurasia and the accretion of Cimmerian terranes to close the 318 PaleoTethys. The vertical sinking velocity of slabs average ~1.6 cm/yr in the mid-mantle between 319 200 and 170 Ma for the Tethyan subduction system, with sinking rates as low as ~1.3 cm/yr in the 320 lowermost mantle between 170 and 150 Ma. We do not estimate upper mantle sinking velocities from our models since slabs are imposed to a depth of 350 km. About 5° of southward advection of 321 322 the unbroken PaleoTethyan slab occurs between 170 and 150 Ma, after which the kinematic 323 boundary conditions on the surface begin to diverge between the conventional and alternative 324 scenarios.

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The conventional scenario has continued Andean-style subduction (Figure 9b), while a large back-326 arc opens in the alternative scenario to shift subduction in the central MesoTethys to ~15°N by 327 328 120 Ma (Figure 9c). A slab window forms between 150 and 120 Ma in both scenarios due to the 329 intersection of the MesoTethyan mid-ocean ridge and the subduction zone [Thorkelson, 1995]. The 330 continuous older Paleo- and Meso- Tethyan slab approaches the CMB, while the subduction of the 331 leading edge of the younger slab is deflected at the transition zone in both scenarios. The younger 332 slab drapes across the transition zone between ~15 and 25°N in the alternative scenario as the 333 subduction zone migrates southward during back-arc opening between 150 and 120 Ma. By 90 Ma,

334 the slab in the conventional model has penetrated the transition zone and is sinking at latitudes of  $\sim$ 25 to 30°N, while the upper mantle slab in the alternative model is only starting to penetrate the 335 336 660 km transition. The lower mantle slab in both models has already started to drape over the CMB, 337 with considerable southward displacement by 90 Ma. Importantly, subduction is occurring at ~15°N 338 in the alternative model by 90 Ma, rather than ~25°N in the Andean-style subduction of the 339 conventional model. The lowermost mantle structure is similar in both models by 60 Ma, while 340 mid- and upper- mantle structure differs considerably in the alternative scenario by 55 Ma due to 341 the initial collision between Greater India and the NeoTethyan island arc. This initial collision 342 causes NeoTethyan slab break-off and initiates the subduction of the back-arc along southern 343 continental Lhasa. Final slab break-off occurs after ~55 Ma in the conventional model following the 344 onset of continental collision and suturing between Greater India and continental Eurasia. The backarc slab detaches by ~40 Ma in the alternative scenario due to the later continental collision 345 346 between Eurasia and a smaller Greater India. Two discrete slab volumes are sinking at mid-mantle depths by 30 Ma in the alternative scenario, in contrast to the single descending slab at similar 347 348 depths in the conventional model. Slab material sourced from the Pacific north of ~35°N at depths 349 greater than ~1200 km appears by 30 Ma and interacts with the Tethyan slabs in the lowermost 350 mantle.

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The present-day prediction of mantle structure from both scenarios (Figure 10a) was compared to 352 equivalent vertical slices in a range of mantle tomographic models (Figure 10b, Appendix E). P-353 354 and S- wave models, suggest the distribution of Tethyan slabs range from near-equatorial latitudes, to as far north as 35°N – with a wide band of coherent positive seismic velocity anomalies at mid 355 mantle depths between ~800 and 2400 km. The numerical model of the conventional scenario 356 357 predicts a single slab at mid-mantle depths (~1200 to 2000 km) with a latitudinal range of ~15 to 25°N, while the alternative scenario predicts two discrete mid-mantle slabs. The back-arc slab is 358 359 found at latitudes between 20 and 25°N at mid-mantle depths, merging with the Paleo- and Meso-

360 Tethvan slabs in the lower mantle, while the youngest NeoTethvan slab is found at depths between 361 ~1000 and 1300 km between ~0 and 15°N. Thus the two numerical model predictions vary the most 362 between ~1000 and 1300 km depth. The mid-mantle slabs that are common to both scenarios at 363 present-day can be accounted for in all seismic tomographic models, albeit the corresponding 364 seismic velocity anomalies are more smeared than the thermal anomaly predicted by the numerical 365 models. All seismic tomography models suggest the existence of a discrete slab, offset further south 366 of the main slab volume observed at mid-mantle depths. The latitudinal offset varies in the 367 tomographic models, with GRAND-S suggesting a small slab between ~10-15°N at ~1500 km 368 depth, while other P- and S- wave models suggest a shallower depth of ~1000 km for this seismic 369 velocity anomaly between ~5°S and 15°N. The conventional scenario does not reproduce the large 370 latitudinal range of slab material in the mid-mantle, while the alternative scenario produces a better 371 match to observations of present-day mantle structure. The continuity of an upper mantle slab with 372 the remaining mid-mantle slab material is questionable as GyPSuM-P and SB4L18 suggest no link 373 across the transition zone, while the remaining tomographic models support this link that is not 374 replicated in either numerical models of Tethyan convergence.

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#### 376 3.2. Lateral distribution of slab material and lateral mantle flow

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378 The horizontal depth slices of seismic tomography from a number of models between ~1000 and 379 1300 km depth highlights the differences between the conventional and alternative subduction 380 scenarios with a large latitudinal range of positive seismic velocity anomalies (Figure 11, 381 Appendix F). At a depth of ~1000 km, the S-wave tomographic models (SAW24B16 and GRAND-382 S) support the existence of slab material further south than can be accounted for by the conventional 383 model (red contour). Although the P-wave MIT-P model also suggests a larger latitudinal range of 384 slab material at this depth, the amplitude of the positive seismic velocity anomaly is diminished in 385 oceanic regions, an expected artefact of most P-wave models. At a depth of ~1250 km, the

386 discontinuity of the main northern slab is better reproduced by the alternative scenario, while both 387 numerical models provide a good match in the Java-Sunda region. The GRAND-S model does 388 suggest the existence of a discrete slab south of the main seismic velocity anomaly, but it is further 389 east than predicted by the alternative subduction scenario. The positive seismic velocity anomalies 390 at this depth extend to equatorial latitudes, which cannot be explained by the conventional scenario 391 of subduction. The match between the southernmost slab predicted by the alternative model at a depth of ~1300 km is highest with SAW24B16 and GyPSuM-P. The amplitude of the 392 393 corresponding positive seismic velocity anomaly in the MIT-P model is diminished, and may be 394 due to the inherent shortcoming of some P-wave models where the amplitude of seismic velocity 395 anomalies diminishes significantly away from continental regions with high seismic ray coverage. 396 The continuity of the main northern slab volume is also unclear at this depth, with SAW24B16 and 397 MIT-P suggesting a discontinuous slab while GyPSuM-P indicates slab continuity. The 1445 km 398 depth slice in GRAND-S shows a large discrete slab south of the main slab volume, but is not 399 reproduced by either the conventional or alternative scenario. This is largely due to the smearing 400 effect in tomographic models, contrasting the well-defined slab contours of thermal anomalies from 401 numerical models.

402

403 The southward extent of our back arc in the alternative scenario was guided by seismic tomography 404 images with the initial assumption of vertical sinking of slab material through the mantle. However, 405 our global flow models show that slabs are advected by lateral return flow at mid-mantle depths, 406 with significant variations in the mantle flow direction and magnitude through time (Figure 12, 407 Appendix G). Our plate motions contain a net rotation component [Torsvik et al., 2010] that induces 408 shear deformation distributed through the mantle. It is beyond the scope of this study to carry out a 409 global analysis of net mantle rotation versus regional lateral mantle flow. For this reason, we limit 410 our analysis of regional lateral mantle flow to its north-south component (red arrows in Figure 12) 411 that is unaffected by longitudinal net rotation (which represents the bulk net rotation in our model).

412 Our results show that the northward subduction of Tethyan slabs leads to a significant southward 413 return flow at mid-mantle depth. A strong southward component of flow dominates at 90 Ma at a 414 depth of ~1000 km, where NeoTethyan slabs are strongly advected. However, the southward component of flow wanes between 60 and 30 Ma, and returns to a strong southward-directed 415 416 advective flow towards the present-day. The advection of slabs reveals that the assumption of 417 vertically sinking slabs used to build a subduction reference-frame [Van Der Meer et al., 2010] is 418 an over-simplification. To the same effect, we also note that the constant sinking rate 1.2 cm/yr in 419 the lower mantle assumed in our initial slab age-coding workflow is slightly smaller than the 1.3 to 420 1.6 cm/yr range evident for the Tethyan region in the numerical models.

- 422 **4. Discussion**
- 423

## 424 4.1. Insights from numerical models

425

Interpretation of slab material in our subduction models relies on the first-order comparisons 426 427 between thermal heterogeneities and seismic velocity anomalies, assuming that high seismic 428 velocity anomalies that correlate across P- and S- wave mantle seismic tomography are colder, 429 denser and therefore represent sinking materials from subduction zones [Simmons et al., 2009]. 430 Although the viscosity contrast between the upper and lower mantle is not well constrained, our 431 results show that a viscosity contrast of 100 between the upper and lower mantle is sufficient to 432 retain coherent Meso- and Neo- Tethyan slabs at mid-mantle depths that match the observed mantle 433 structure from seismic tomographic models and are consistent with previous studies [Hager, 1984; 434 Jarvis and Lowman, 2005]. Our global subduction models show that the conventional scenario of long-lived Andean subduction along continental Eurasia does not reproduce the latitudinal range of 435 436 slab material at mid-mantle depths at present day. Our models also show that the subduction of additional oceanic lithosphere, such as a large back-arc basin, is necessary to account for the 437 438 present-day mantle structure, and that the overall chronology likely involved initial contact between 439 Greater India and an intra-oceanic island arc.

440

The alternative scenario in our study reproduces discrete parallel slab volumes at mid-mantle depths, largely consistent with interpretations of *Hafkenscheid et al.* [2006] and *Van der Voo et al.* [1999b]. While most seismic tomography models at depths of ~1000 to 1300 km suggest a southwardly-offset slab, the longitudinal position of this volume varies somewhat across the tomographic models. Although our numerical models predict a significant westward mantle flow at mid-mantle depths (Figure 12, Appendix G), we do not interpret this observation, as our imposed plate kinematics imply a ~0.12°/Myr westward net rotation of the lithosphere as described by 448 Torsvik et al. [2010] and therefore it is difficult to isolate the local east-west component of flow 449 from the effects of global lithospheric net rotation. Independent of longitudinal flow, our models 450 suggest a southward flow of Tethyan slab material in the mantle that we interpret as return flow 451 from north-directed subduction. Therefore, we suggest that while it is useful to the first order, the 452 assumption of vertically sinking slabs and constant sinking rates [Van Der Meer et al., 2010] is an oversimplification of complex mantle flow. An improvement to our method of age-coding slabs in 453 454 tomography would be to backward-advect slabs from seismic tomography using an adjoint 455 convection model in order to correct for lateral mantle flow as demonstrated by Liu and 456 Gurnis [2008].

457

The longitudinal location of the back-arc slab may also vary if we take into account the evolution of 458 459 the Western and Easternmost Tethys, which was beyond the scope of this study. The motion of 460 Argoland (West Burma Block) in our models, as a Gondwana-derived terrane that accreted onto Eurasia at ~80 Ma [Heine and Müller, 2005], may have closed a back-arc that extended along the 461 462 entire Eurasian margin as in the model proposed by Stampfli and Borel [2002]. In addition, there 463 may have been one or more episodes of back-arc opening in the Tethys as observed in the western 464 Pacific [Clark et al., 2008]. In our alternative scenario we implement a long-lived Tethyan back-arc 465 that is in some ways similar to the Kohistan-Ladakh back-arc system, thought to have opened at 134  $\pm$  3 Ma based on K-Ar ages and closed well before the India-Eurasia collision [*Allegre et al.*, 1984; 466 Khan et al., 2007; Pudsey, 1986]. However, some authors have suggested that India first collided 467 468 with the Kohistan-Ladakh magmatic arc by 55 Ma [Gaetani and Garzanti, 1991; Khan et al., 2009]. 469 Earlier studies, such as Dewey et al. [1988], had also suggested a late continental collision at ~45 Ma, with a Cretaceous-age emplacement of the Spongtang Ophiolite (Figure 1). However, 470 471 there is little evidence that the Kohistan-Ladakh or the Spongtang arc system extended eastward 472 along the Indus-Tsangpo suture as in our alternative scenario. If the initial slowdown of India at 473 ~55 Ma was unrelated to collision with an island arc, then the subduction of a large buoyant 474 structure on the downgoing plate, such as an oceanic plateau or seamount chain, may have impeded 475 subduction and caused a slowdown in convergence. However, we predict that such a scenario 476 would not result in a large latitudinal range of slab material as observed in mantle tomography. 477 Apart from the subduction of a back-arc basin, it is also possible to subduct additional oceanic 478 material along southern Eurasia if a microcontinent had rifted off Greater India to create a seaway 479 as proposed by Roy [1976]. However, little trace of such features exists in the suture zone and so 480 they were either completely subducted, not yet mapped in the remote regions of the collision zone 481 or they did not exist.

482

483 In addition to the possibility of intra-oceanic subduction significantly altering the dynamics of India-Eurasia convergence, mantle plumes have recently been implicated in accelerating India's 484 485 advance toward Eurasia prior to collision [Cande and Stegman, 2011; van Hinsbergen et al., 486 2011a]. Cande and Stegman [2011] associate the acceleration of India at the onset of Anomaly 30 (~67 Ma) with the arrival of the Reunion plume head at the India-Africa ridge and to the 487 488 emplacement of the voluminous Deccan Traps on the Indian continent. They find that the motion of 489 India slows considerably by 52 Ma, followed by a more significant decrease in speed after 45 Ma 490 accompanied by a change in spreading direction between India and Africa. A scenario with initial 491 collision at ~52 Ma would require a smaller back-arc basin than we implemented or a smaller 492 Greater India, leading to a later initial collision and possible continental contact at ~45 Ma. This is 493 plausible as both the extent of Greater India and the proposed NeoTethyan back-arc basin are 494 uncertain.

495

#### 496 4.2. Insights from surface geology

497

498 The large back-arc basin of the alternative scenario better fits observations of subducted slab 499 material in mantle tomography to suggests that pre-Eocene subduction occurred at least 10° further

500 south than can be explained by subduction at an Andean-style margin proposed by conventional 501 models. To produce Andean-style subduction this far south would imply that either Lhasa was 502 much further south at this time or that the entire Asian continent was decoupled from European 503 continent. However, the potential regions of decoupling between Europe and Asia, notably the Ural 504 mountains, have been discounted as the source for the large-scale independent motion of Asia relative to Europe in the Cenozoic [Lippert et al., 2011]. A recent review of paleomagnetic data 505 506 from the volcanic Linzizong Group by Yi et al. [2011] suggests that Lhasa was at a latitude of  $6.1 \pm$ 507  $8.5^{\circ}N$  (64-60 Ma),  $12.9 \pm 4.6^{\circ}N$  (60-50 Ma) and  $19.3 \pm 4.7^{\circ}N$  (50-44 Ma) in the early Cenozoic. 508 The study by Sun et al. [2012] indicates that the Lhasa terrane was at  $15.2 \pm 6.3^{\circ}$ N at collision 509 timing of ~54-47 Ma derived from the overlap of Tethyan Himalaya apparent polar wander paths 510 with those of Lhasa. Others report that the paleolatitude of the Lhasa terrane was not further south 511 than 20°N throughout Eocene time based on paleomagnetism of Paleogene volcanics [Lippert et al., 512 2011]. Unfortunately, paleomagnetic data alone places Lhasa anywhere between being south of the equator or in excess of 20°N in pre-collision times. However, equatorial or low latitude positions of 513 514 Lhasa in the early Eocene would imply extreme post-collisional north-south shortening of the 515 terrane, where only ~290 km north-south shortening of the terrane is recorded in the geology from ~100 to 20 Ma [van Hinsbergen et al., 2011b]. 516

517

In the region of Mt Everest, an initial high pressure and temperature metamorphic event at ~39 Ma 518 519 [Cottle et al., 2009] may be indicative of continent-continent collision between Greater India and 520 Lhasa. This would explain a further significant drop in India-Eurasia convergence rates by ~40 Ma (Figure 2). Additionally, studies of sediments near Mt Everest indicate a marine depositional setting 521 prevailing as late as 34 Ma [Wang et al., 2002], suggesting that suturing was likely diachronous 522 523 along the margin and that final collision between India and Eurasia occurred much later than India's 524 initial slowdown. Reactivation of the major Altyn Tagh fault by 40 Ma [Liu et al., 2007], strike-slip motion along the Ailao Shan-Red River shear zone by 35 Ma [Leloup et al., 2007; Leloup et al., 525

526 2001] and post-collisional deformation propagating north with shortening and exhumation 527 propagating north in the Tian Shan at ~25 Ma [Dumitru et al., 2001] suggest an early Paleogene continent-continent collision would require a ~20 Myr time lag in regional geological responses. 528 529 The absence of volcanic activity along southern Lhasa, at least in the late Cretaceous [Chung et al., 530 2005], also casts doubt on continuous Andean-style subduction in the NeoTethys. However, the 531 first pulse of Transhimalayan Batholith emplacement at ~100 Ma [Debon et al., 1986; Miller et al., 532 1999] is incompatible with our alternative scenario, and may suggest short-lived pulses back-arc 533 seafloor subduction. The Linzizong Volcanics (Figure 1), previously interpreted as Andean-style 534 subduction-related suites, along with another pulse of Transhimalayan Batholith emplacement 535 [Debon et al., 1986; Miller et al., 1999], between ~65 and 40 Ma [Xia et al., 2011] are observations that are compatible with the closure of a back-arc basin in our alternative scenario. Similarly, the 536 generalized stratigraphic interpretations of Aitchison et al. [2011] indicate a hiatus in pluton 537 538 emplacement from ~80 to 65 Ma. However, the duration of that magmatic gap depends on how well 539 temporally- and spatially- sampled the magmatic suites are in the collision zone. We speculate that the Linzizong Volcanics are emplaced as a result of Andean-style subduction of the NeoTethyan 540 541 back-arc between ~65 and 40 Ma, after which continental deformation and extrusion dominated.

542

## 544 **5.** Conclusion

545

546 Our study demonstrates that competing end-member kinematic scenarios of plate convergence can 547 be tested using global subduction models whose predictions can be validated using both surface 548 geology and sub-surface mantle structure. Numerical models of NeoTethyan intra-oceanic subduction better reproduce the volumetric, vertical and lateral distribution of slab material 549 550 interpreted from P- and S- wave mantle tomography models than long-lived Andean subduction 551 along southern Lhasa. In this way we are able to show that the NeoTethys was likely consumed 552 along an intra-oceanic island arc that was accreted to the leading margin of Greater India during the 553 initial convergence rate drop with Eurasia at ~60 Ma. The subduction of additional oceanic crust 554 from a Cretaceous-age back-arc along southern Lhasa is also largely consistent with the timing of 555 magmatic episodes north of the Yarlung-Tsangpo Suture Zone and the large latitudinal range of 556 positive seismic velocity anomalies in seismic tomography beneath the convergence zone. The significant convergence rate drop between India and Eurasia by ~40 Ma correlates to the final 557 558 continent-continent collision that better explains changes in magmatism [Debon et al., 1986; Miller 559 et al., 1999; Xia et al., 2011], high pressure/temperature metamorphism near the suture zone [Cottle et al., 2009], marine deposition as young as ~34 Ma [Wang et al., 2002] and the onset of Indochina 560 extrusion by ~35 Ma [Leloup et al., 2007; Leloup et al., 1995; Tapponnier et al., 1990]. In contrast, 561 562 a continent-continent collision soon after ~60 Ma as proposed by the conventional models requires 563 extreme shortening of Lhasa, an extremely large Greater India and a 20 Myr time lag to account for 564 these major events recorded in the regional geology.

## 566 Acknowledgements

567

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576

578 Figure Captions

579

Figure 1. Regional geology related to the India-Eurasia collision, showing Tethyan suture zones,
magmatic episodes and major faults [*McDermid et al.*, 2002; *Styron et al.*, 2010].

582

**Figure 2.** India-Eurasia convergence rates derived at 1 Myr intervals from the motion of a presentday reference point (30°N, 80°E) using alternative rotation models that were standardised to the time scale of *Cande and Kent* [1995]. The initial drop in convergence rates varies from ~58 Ma in the *Lee and Lawver* (1995) model, to ~58 and ~48 Ma in the model of *Molnar and Stock* (2009), while the model of *Müller et al.* (2008) suggests initial convergence rate drop at ~55 Ma. Thus the interpretation of collision timing purely from convergence rates is largely model-dependent, with a 10 Myr range of initial collision timing just from four rotation models.

590

Figure 3. Timing of contact implied by different rotation models with the conventional pre-591 592 collision geometries of Greater India [Lee and Lawver, 1995] and Andean-style Eurasian margin. A 593 number of rotation models were implemented for India, standardised to a common time scale 594 [Cande and Kent, 1995] and with a fixed Eurasia. A number of other subduction zone paleo-595 locations are plotted for comparison, highlighting that the choice of pre-collision margins has a 596 more significant effect on collision timing than the choice of rotation models alone. The choice of 597 Eurasian margin geometries result in ~15 Myr timing difference of the initial contact between 598 Greater India and Eurasia whereas the different rotation models can only account for ~5 Myr 599 difference in interpreting initial collision based on an Andean-style convergence history for the 600 NeoTethys. The collision between Greater India and Lhasa (green) occurs at ~55 Ma in the Lee and 601 Lawver (1995) model. We chose to largely follow the subduction zone location as proposed by 602 Replumaz et al. (2004) for pre-collision times in the conventional Andean-style margin, with the

603 Lee and Lawver (1995) extent of Greater India resulting in continental contact at ~58 Ma. Light
 604 grey shading represents extent of Eurasian continental crust in our model.

605

**Figure 4.** Pre-breakup Gondwana fit (fixed Australia) at 140 Ma with present-day *Sandwell and Smith* (1997) 1-min gravity anomaly grid. The conventional model (*Lee and Lawver* [1995]) proposes a large hypothetical extent of Greater India [black line]. The alternative model invokes a smaller Greater India [orange line] limited in extent by the Gondwana pre-breakup fit and the Wallaby Zenith Fracture Zone (WZFZ) [white line]. India rotated counter-clockwise away from Australia, followed by a northward advance that is highlighted by the Indian Ocean fracture zone bends.

613

614 Figure 5. End-member kinematic models, conventional model [left] and alternative model [right], 615 with plate velocities, block outlines, plate boundaries and seafloor age-grids. The conventional model implies northward dipping Andean-style subduction along southern Eurasia to consume the 616 617 Paleo-, Meso- and Neo- Tethys. The alternative model invokes back-arc spreading from 150 to 618 120 Ma in the central MesoTethys. Argoland (West Burma Block) accretes to Eurasia by 80 Ma in 619 both scenarios. Greater India collides with Eurasia by 55 Ma in the conventional model to terminate 620 NeoTethyan subduction. The NeoTethyan island arc is accreted onto a smaller Greater India by 621 ~58 Ma in the alternative model, after which the associated back-arc is subducted along an Andean-622 style margin until continental collision between Indian and Eurasia by ~40 Ma.

623

**Figure 6.** Effect of absolute reference frames on past locations of subduction in the Tethys and northwest Pacific. Present-day coastlines are plotted only for reference. Latitudinal position of Tethyan subduction varies by up to 10° latitude across different absolute reference frame at ~200 Ma. From 150 Ma, moving hotspot and true polar wander-corrected reference frame (MHS/TPW) is similar to the fixed hotspot reference (FHS), with the subduction reference frame
(VMSR) differing largely in longitudinal position. [Orthographic projection, centred: 0°N, 90°E]

631 Figure 7. Plate reconstructions with age-coded mantle seismic tomography (MIT-P and GRAND-S), applying an average slab sinking rate of 3 and 1.2 cm/yr in the upper and lower mantle 632 633 respectively, as an independent constraint on the southward extent of the NeoTethyan back-arc 634 basin at pre-collision times. Location of northward-dipping subduction along continental Eurasia in 635 the conventional scenario (left column) does not account for the large positive seismic velocity 636 anomaly at pre-collision times, while a subduction zone offset further south in the central Meso-637 and Neo- Tethys in the alternative scenario (right column) better describes the P- and S- wave 638 tomographic observations, assuming near-vertical sinking slabs. Reconstructed Coastlines/Block 639 outlines - grey, GI - Greater India, MT - MesoTethys, NT - NeoTethys. [Orthographic projection, 640 centred: 0°N, 90°E]

641

Figure 8. Model horizontal average non-dimensional temperature (left) and viscosity (right) withdepth for the initial condition (red) and present-day (black).

644

645 Figure 9. Time-dependent horizontal slices (328 km) of predicted temperature field with plate 646 boundary configurations and the location of vertical slices (orange - static great circle segment between 10°S, 55°E and 45°N, 90°E) documenting mantle evolution resulting from Tethyan plate 647 648 motions. a) Initial condition between 200 and 150 Ma depicting the accretion of Cimmerian terranes 649 of Iran, Afghanistan, South Tibet (Lhasa) and Sibumasu onto the Eurasian margin in the late Jurassic with Andean-style subduction at ~30°N. b) Conventional models of Meso- and Neo-650 651 Tethyan evolution have continuous Andean-style subduction, with collision between India and Eurasia at ~55 Ma where subduction of oceanic lithosphere ceases and a single slab descends 652 653 towards the CMB at ~25°N. c) Equivalent vertical slice through the alternative scenario highlights

654 the different location of subduction in the Meso- and Neo- Tethys during the opening of the back-655 arc between 150 and 120 Ma, shifting subduction south to ~10°N between 120 and 60 Ma. Slab 656 penetration is slowed at the upper/lower mantle transition zone, with a mantle avalanche occurring 657 after 90 Ma. A smaller Greater India collides with the island arc to consume the back-arc basin along continental Eurasia (southern Lhasa) between ~55 and 40 Ma, resulting in two discrete slab 658 659 volumes at ~10°N and ~25°N at mid-mantle depths by 30 Ma. PaleoTethyan material subducted 660 originally at ~30°N in both scenarios migrates gradually southward due to the return flow at mid-661 and lower-mantle depths.

Figure 10. a) Present-day temperature field predicted by numerical models along a vertical profile 662 663 (magenta line - location of vertical, great circle segment between 10°S, 55°E and 45°N, 90°E), with 664 slabs defined as thermal anomalies colder than 10% of the ambient mantle temperature. b) Present-665 day mantle structure from a range of P- and S-wave mantle tomographic models along equivalent vertical slices, with slab contours from numerical models of the conventional (red) and alternative 666 (green) models of India-Eurasia collision. The conventional model predictions do not account for 667 668 the large latitudinal range of seismic velocity anomalies at mid-mantle depths. Tomographic models suggest the existence of a slab at ~1000 km depth at near-equatorial latitudes, and a larger mid-669 670 mantle slab ranging from 5 to 35°N. The conventional scenario only reproduces the larger slab, and 671 does not account for the large latitudinal range of seismic velocity anomalies that are better 672 reproduced by the alternative scenario invoking intra-oceanic subduction in the Meso- and Neo-673 Tethys.

674

Figure 11. Comparison of depth slices from P- and S- wave models with slab contours at similar depths from conventional (red) and alternative (green) subduction model outputs. Present-day coastlines are plotted for reference (dark grey). Positive seismic velocity anomalies better correlate with the alternative scenario slab contours to reproduce the discrete slab volumes and the large latitudinal range. At depths of ~1200 and ~1300 km, the discrete slab volume predicted by the

680	alternative scenario is also observable in tomography, with a systematic $\sim 10^{\circ}$ longitudinal offset in
681	some tomographic models. A depth slice at ~1300 km was not available for the GRAND-S model,
682	and so the closest slice of 1445 km was included. [Mercator projection]

683

**Figure 12.** Time-dependent mantle flow predicted by the conventional (left) and alternative (right) scenarios of India-Eurasia convergence at 30 Myr increments between 90 and 0 Ma at a depth of 1012 km. Predicted lateral flow is denoted by grey arrows, and the north-south component is highlighted by larger red arrows. A strong southward component in the return flow at this depth is predicted at 90 and 0 Ma. [Albers projection, centre: 25°N, 80°E, standard parallels: 10 and 40°N]

691 Tables

692

Table 1. Collision timing between India and an Andean-style margin of Eurasia as inferred from the contact of maximum extent Greater India (*Lee and Lawver*, 1995) and a range of southern Eurasian margins using different rotation models that were standardised to the time scale of *Cande and Kent* (1995). Timing offset is a maximum of 4 Myr for each proposed margin, but 15 Myr when comparing all margins – showing that the pre-collision margin choice has the biggest impact on the interpretation of collision timing rather than the choice of motion models of India relative to Eurasia.

700

		South Eurasian Andean-style margin			
		Tapponnier	Lee and	Norton	Replumaz et
		et al. (1986)	Lawver	(1999)	al. (2004)
			(1995)		
	Müller et al. (2008)	61 Ma	53 Ma	53 Ma	55 Ma
	– this study				
del	van Hinsbergen et	61 Ma	53 Ma	53 Ma	55 Ma
Mo	al. (2011)				
ation	Molnar and Stock	60 Ma	49 Ma	50 Ma	53 Ma
Rota	(2009)				
	Lee and Lawver	64 Ma	54 Ma	53 Ma	57 Ma
	(1995)				

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702

# 703 **Table 2.** Vertical and lateral mesh resolution in *CitcomS* models

		Average Resolution
Vertical	Upper mantle	26 km
	Lower mantle	56 km
Lateral	Surface	50 km
	Core-mantle boundary	23 km

# **Table 3.** Numerical model parameters

Constant variables				
Reference density, $\rho_0$	4000 kg/m <sup>3</sup>			
Reference viscosity, $\eta_0$	$1 \times 10^{21}$ Pa s			
Activation energy (upper mantle), $E_{\eta}$	100 kJ/mol			
Activation energy (lower mantle), $E_{\eta}$	33 kJ/mol			
Activation temperature, $T_{n}$	225 K			
Thermal diffusivity, κ	$1 \times 10^{-6} \text{ m}^2/\text{s}$			
Coefficient of thermal expansion, $\alpha$	$3 \times 10^{-5} \text{ K}^{-1}$			
Earth radius, <i>R</i>	6371 km			
Gravitational acceleration, g	9.81 m/s <sup>2</sup>			
Temperature contrast, $\Delta T$	1400 K			

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Time (Ma)

Convergence Rate (mm/yr)























Temperature (Non-dim)