Incorporating Raster and Vector Constraints in Plate Reconstruction Modelling

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WARNING:

The first time you import an age-coded raster, GPlates will take time to create some cache files (this can take 5 or more minutes).

The cache files that GPlates creates in the same folder are quite large (up to 100 Mb each), meaning that you need to have enough storage space.

Aim

This tutorial is designed to introduce the reader to the available GPlates rasters, and how to use them, in combination with reconstruction tools such as flowlines and motion paths, to explore possible tectonic histories.

Included Files

[Click here](#) to download the data bundle for this tutorial.
The tutorial dataset (3.4-Incorporating_Raster_and_Vector_Constraints_in_Plate_Reconstruction_Modelling.zip) includes the following files:

GPlates project file: 3.4-Incorporating_Raster_and_Vector_Constraints.gproj

Free Air Gravity Raster files:
- Free_Air_Gravity_Anomalies_Legend.jpg
- Free_Air_Gravity_Anomalies.gpml
- Free_Air_Gravity_Anomalies.jpg

Global Geology Raster files:
- Global_Geology_Legend1.png
- Global_Geology_Legend2.png
- Global_Geology.gpml
- Global_Geology.png

Feature Collections Directory (loaded in the project file by default):
- Rotation File:
  - Matthews_etal_GPC_2016_410-0Ma_GK07.rot
- Static Polygons:
  - Muller_etal_AREPS_2016_StaticPolygons.gpmlz
- Plate Topologies:
  - Matthews_etal_GPC_2016_MesozoicCenozoic_PlateTopologies.gpmlz
  - Matthews_etal_GPC_2016_Paleozoic_PlateTopologies.gpmlz
  - Matthews_etal_GPC_2016_TopologyBuildingBlocks.gpmlz
- Rasters:
  - Seafloor_Age_Grid.gpml
  - Topography.gpml
- Coastlines/Continent-Ocean Boundaries:
  - Matthews_etal_GPC_2016_Coastlines.gpmlz
  - Muller_etal_AREPS_2016_ContinentOceanBoundaries.gpmlz
  - AREPS_GreaterIndia.gpml
- Ridges:
  - Muller_etal_AREPS_2016_Ridges.gpmlz

This tutorial dataset is compatible with GPlates 2.0 and newer.

Background

The GPlates sample data provides users with a number of colour, numerical and time-dependent rasters (details in the Appendix), which display present day features of the crust, from crustal strain, to free air gravity anomalies. As explained in previous tutorials (3.1 and 3.2), these can be 'cookie-cut' according to present day plate geometries, and reconstructed back through time.

In this tutorial we will be investigating the evolution of the Indian Ocean, which began formation during the Cretaceous, following the break up of East Gondwana (Gibbons, 2013).

Before we begin visualizing and exploring some of the available rasters, it may be helpful to understand the difference between the three types of raster files:

- Colour raster – a simple 2-dimensional grid of pixels stored as JPEGS or grid files. The majority of the rasters come in this format.
- Numerical raster - this is similar to a colour raster, however each pixel/cell has a value assigned to it. These values are used to generate the raster, and as such, the colour scale can be changed, unlike colour rasters.
- Time dependent raster –a series of rasters, similar to colour rasters, which have been age-coded so we can observe the evolution of a dataset through time.

Exercise 1 – Evolution of the 90 East Ridge

The 90 East Ridge is a linear, age-progressive plume trace, located in the eastern Indian Ocean, and as its name suggests, strikes almost parallel with the 90° E meridian (Figure 1). It has been referred to as a 'leaky transform', perhaps resulting from the interaction of a transform fault with the Kerguelen plume, which first erupted 117 Ma forming the Rajmahal Traps (O'Neill 2005). The ridge therefore displays an age progression, but not one as simple as an intra-plate chain like the Hawaiian-Emperor volcanic chain displays. The Kerguelen plume appears to have a complicated history and remains active today, driving the volcanism we see on Heard and McDonald Islands located on the central Kerguelen Plateau (O'Neill 2005).

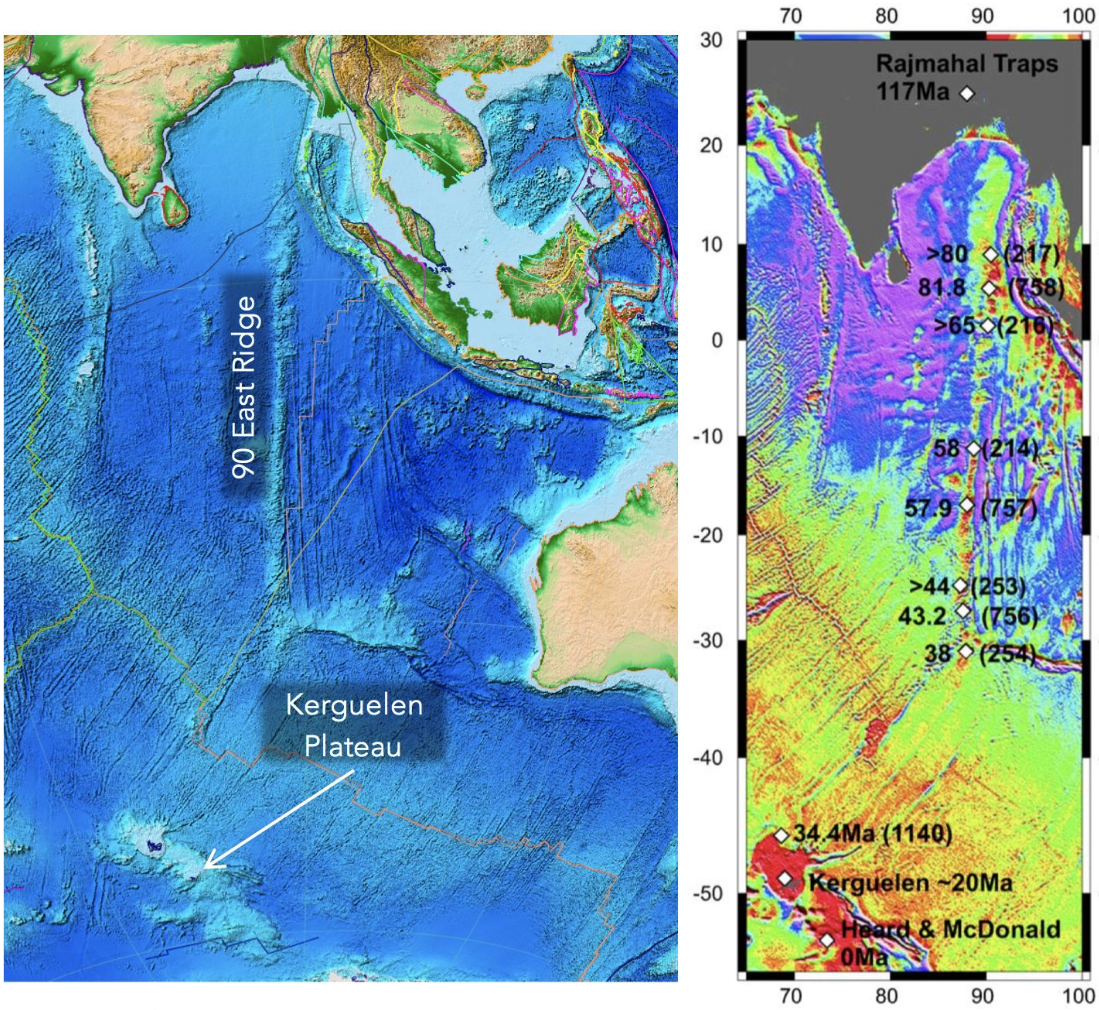


Figure 1. Left: The Indian Ocean as depicted by the 'Topography' raster, indicating the locations of the 90 East Ridge and Kerguelen Plateau. Right: Gravity map of the Kerguelen hot spot system indicating the age progression across the Plateau and Ridge (from O'Neill et al., 2005).

In this exercise we will visualize the 90 East Ridge using the 'free air gravity anomalies' and 'global geology' rasters, and trace its evolution using both flowlines and motion paths. This will highlight the difference between absolute and relative motions and the nature of seafloor fabrics.

(Note: you may notice in some of the figures that the polylines representing the motion path, flowline or plate boundaries appear thicker than they will on your GPlates globe. This has been done to aid with identification. For an explanation of how to do this, refer to Exercise 2)

1. Open GPlates and load in the file 'Tutorial_3.4.gproj' using File > Open Feature Collection or by dragging and dropping the file onto the globe.

2. Rotate the globe so that the Indian Ocean is in view (Figure 2).

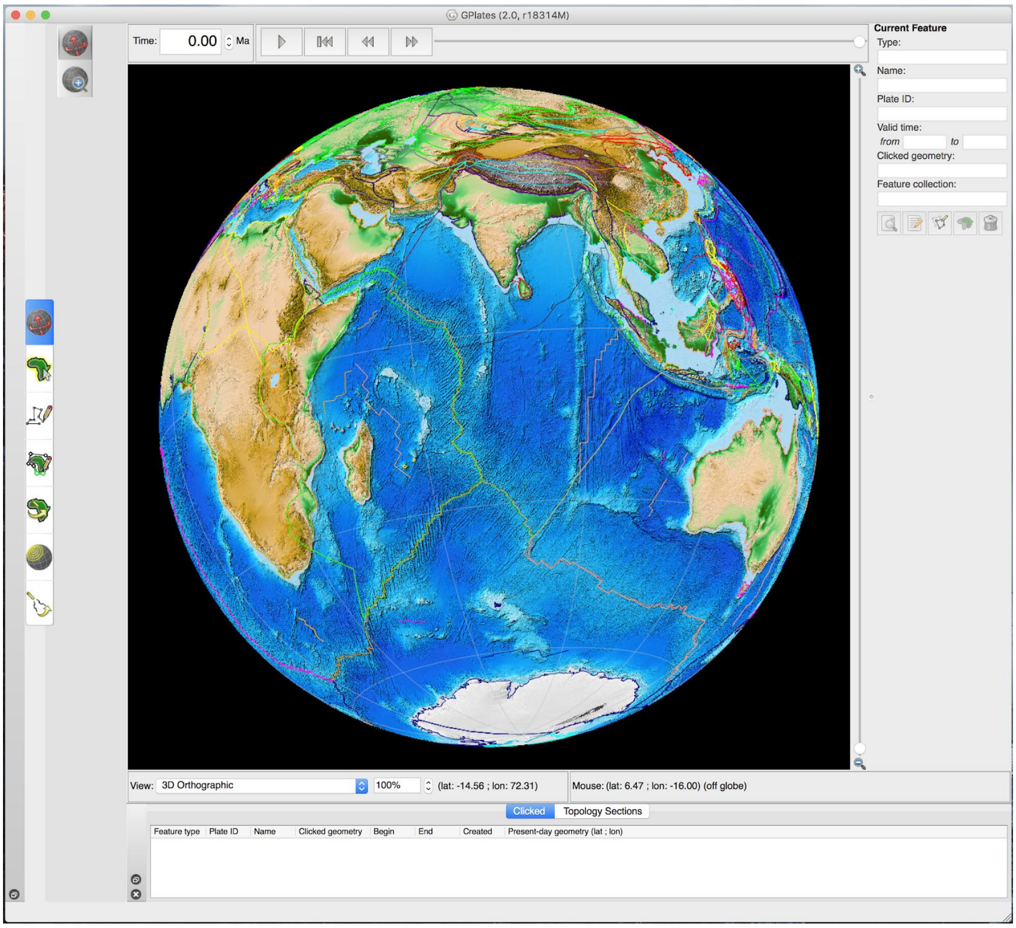


Figure 2. Orientation of globe with Indian Ocean in view

3. If not already visible, bring up the Layers window (Window > Show Layers).

Notice that the 'Topography' raster is visible (as indicated by the  symbol). This layer is red, indicating that it is a raster layer.

See if you can locate the 90 East Ridge and the Kerguelen Plateau. You can also try reconstructing back in time by either entering a time into the dialog box or using the slider and arrows. Can you get an idea of how the 90 East Ridge formed in relation to plate boundaries?

4. Now locate and load the raster file 'Free_Air_Gravity_Anomalies.gpml' by dragging and dropping the file onto the globe or load the .jpg with the same name using File > Import > Import Raster… and answer yes to 'importing the existing GPML file. Your globe should now look like the one in Figure 3.

This provides us with another way of visualizing the crust, in this case, using free air gravity anomalies. Can you still locate the 90 East Ridge? The relative thickness of the ridge compared to the surrounding ocean crust gives it its positive gravity anomaly, which allows us to identify it.

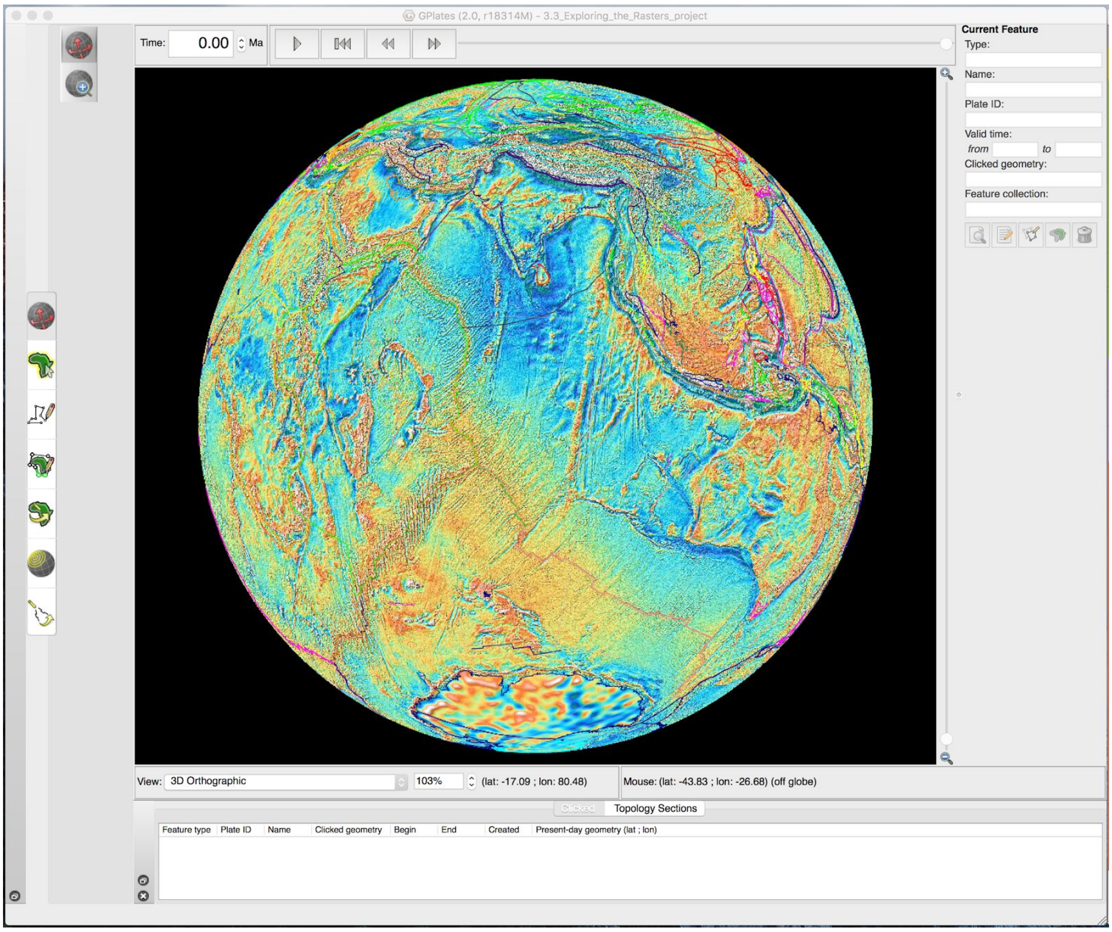


Figure 3. Globe with free air gravity raster visible

5. Expand the red ‘Free_Air_Gravity_Anomalies’ layer in the Layers window by clicking the little black triangle on the left.

6. In the “Inputs” section of the layer, click on the “Add new connection” button under “Reconstructed Polygons:” and select the static polygons file from the list. Under “Age grid raster:” again click on the “Add new connection” button and select “Seafloor_Age_Grid”. Figure 4 displays how your layers window should look after making these changes.

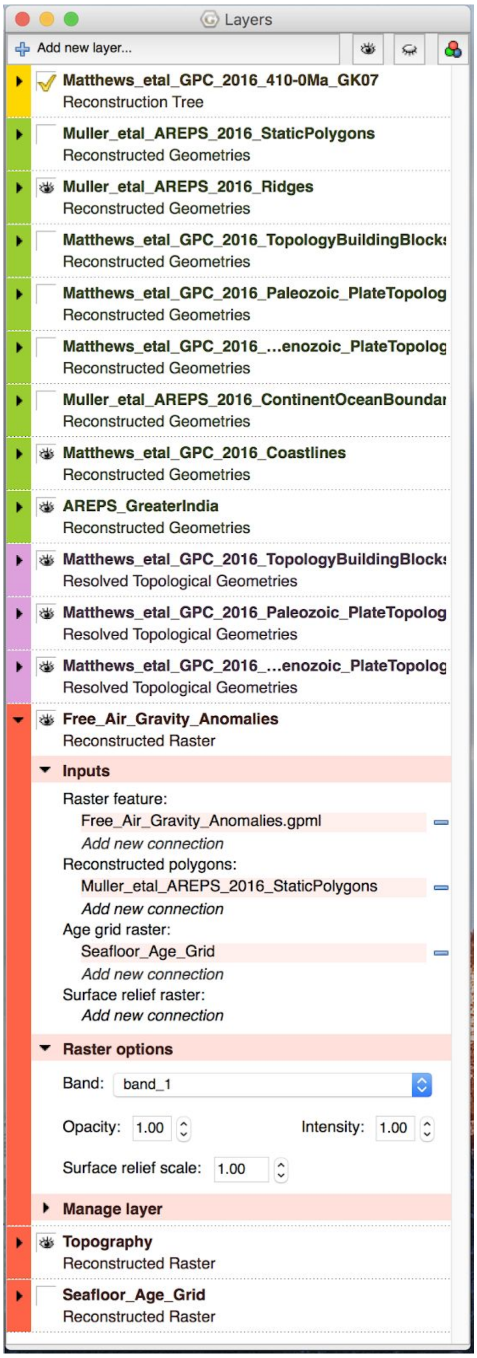


Figure 4. Layers window displaying changes made in step 6

This raster has now been cookie-cut according to the plate polygons, allowing us to reconstruct it back in time. Use the slider to see for yourself. For more information on how this is done refer to Tutorial 3.2: Rotating Rasters and Age-Based Masking of Raster Data. Note that you can continue this tutorial with whichever raster you prefer displayed, but while constructing the flowline and motion path, this tutorial will have the topography raster visible.

7. Hide the 'Free_Air_Gravity_Anomalies' raster so that the 'Topography' raster below it is visible (Figure 5).

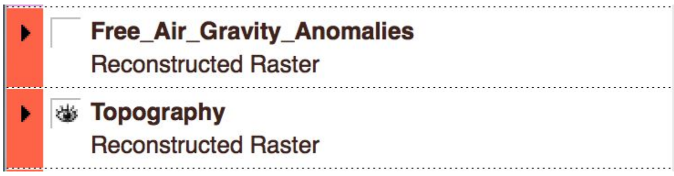


Figure 5. Layers window displaying changes made in step 7

It's now time to construct a flowline and a motion path to track the evolution of the 90 East Ridge. For detailed explanations of these tools refer to Tutorial 2.3: Flowlines and Motion Paths.

Let's begin with the flowline. Flowlines are half stage rotations calculated by GPlates based on the rotation file you are using. They track the relative plate motion away from spreading ridges. In the real world, these develop as features such as fracture zones.

8. Zoom in to focus on the Indian Ocean (Figure 6).

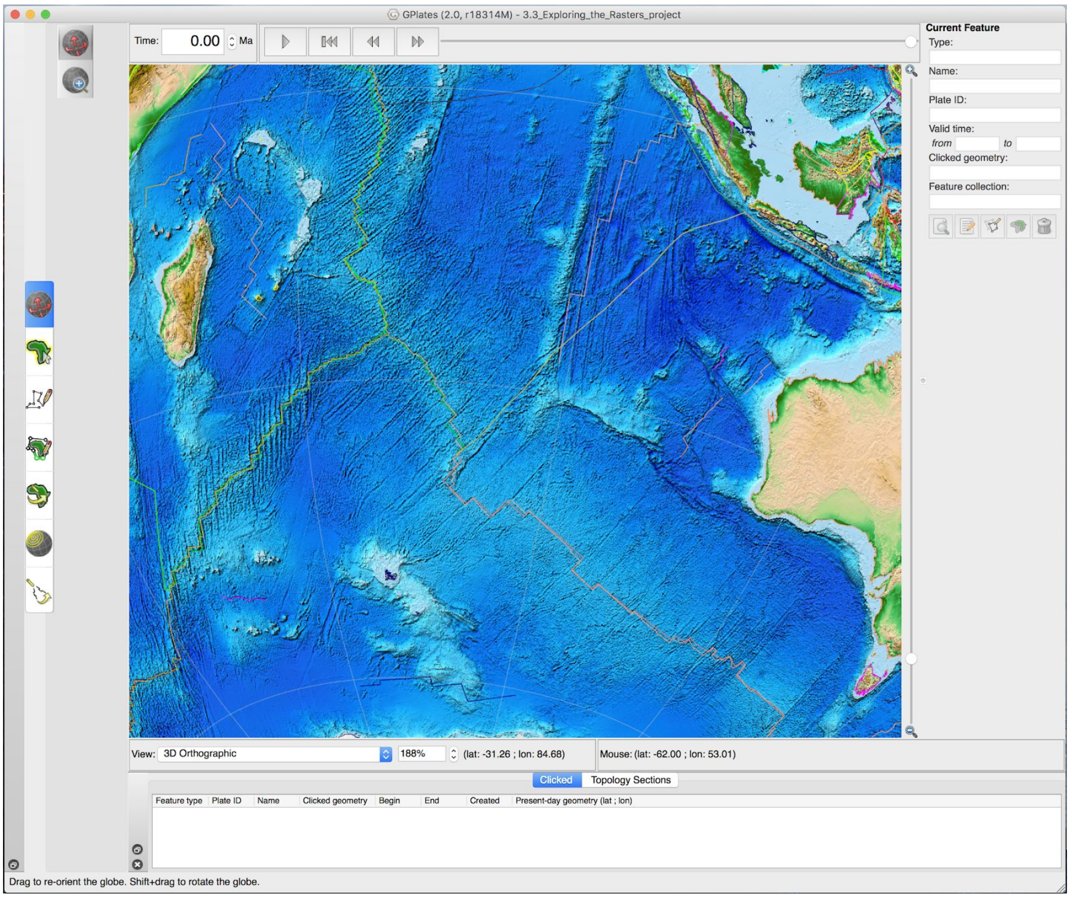


Figure 6. View of the Indian Ocean, Kerguelen Plateau and 90 East Ridge.

9. Select the Digitisation  workflow tab and the Digitise New Multi-point Geometry  tool from its submenu. Use this to create a new seed point located on the spreading ridge between the southern tip of the ridge and the plateau (refer to Figure 7).

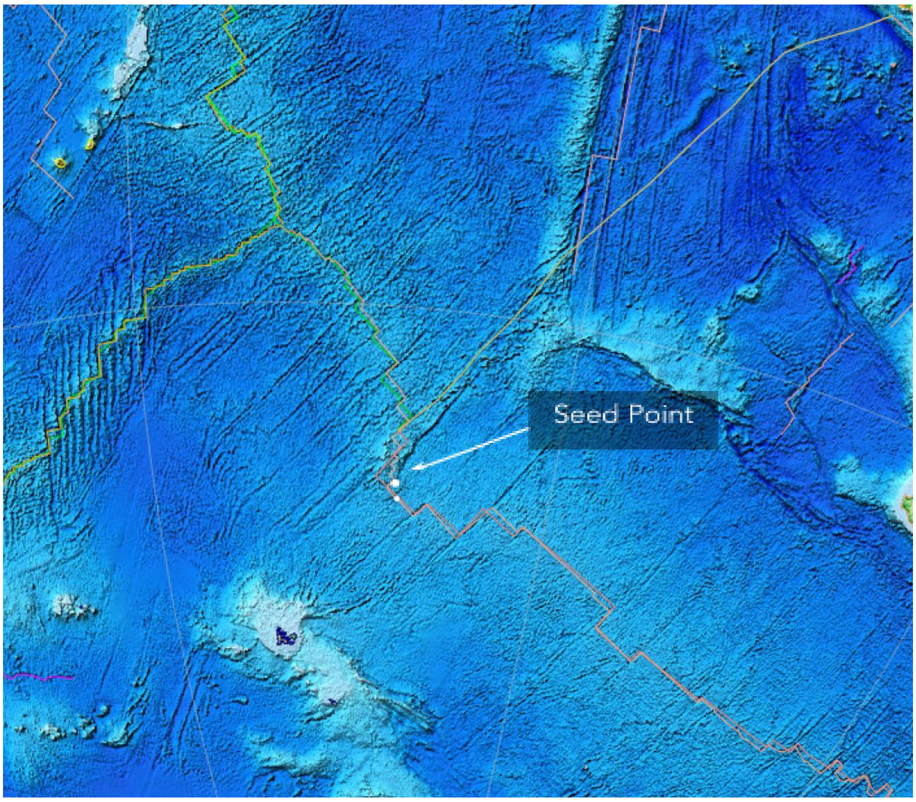


Figure 7. Location of the seed point you will use to construct the flowline.

10. Once you are happy with the location of your point, click on the "Create Feature…" button on the right side of the globe. This will open up the Create Feature menu.

11. Choose your "Feature Type" to be 'Flowline' (Figure 8) and click "Next".

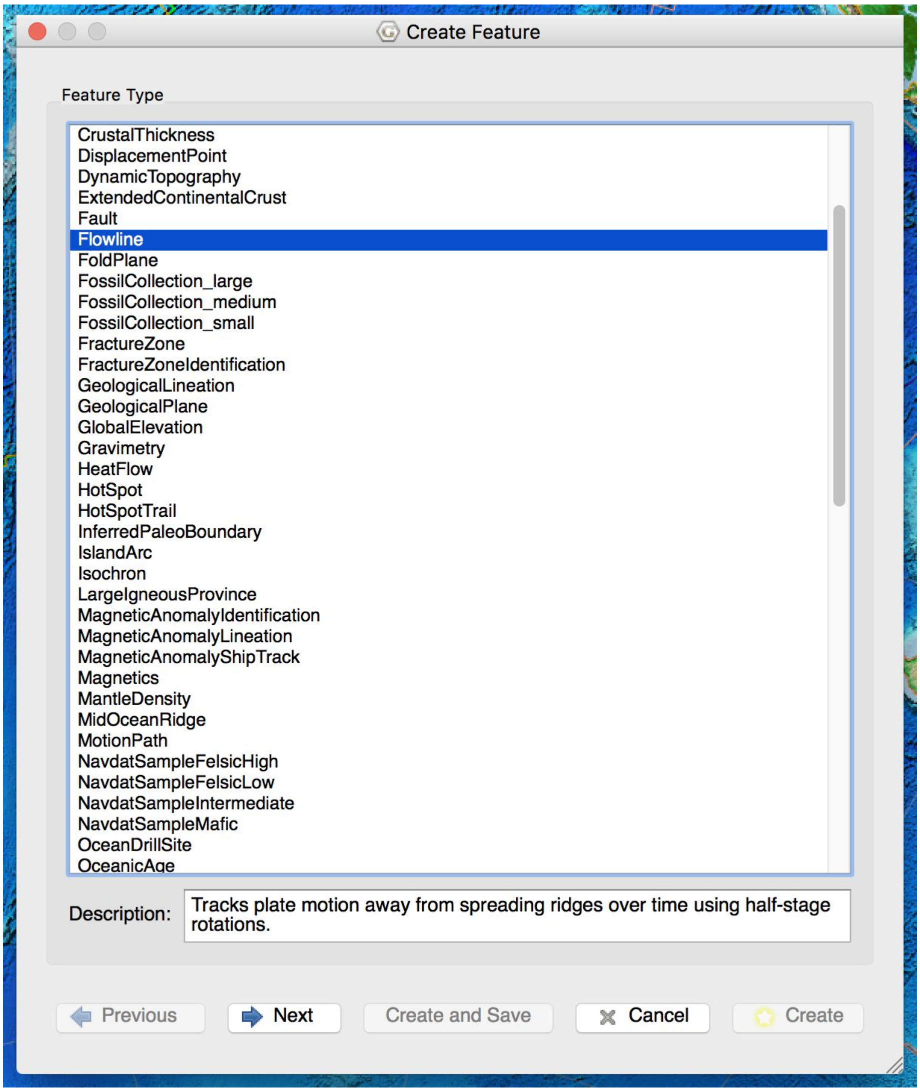


Figure 8. Create Feature menu with 'Flowline' highlighted.

12. This window allows you to fill in the properties of your point. Leave the 'Interpret provided geometries' option as Spreading centre(s). Under 'Common Properties', fill in the following fields (Figure 9):

Left Plate ID: 511 (Central Indian Basin)
Right Plate ID: 802 (Antarctica)
Begin (time of appearance): 83 Ma
End (time of disappearance): select 'Distant Future'
Name: 90 East Ridge Flowline

Click "Next".

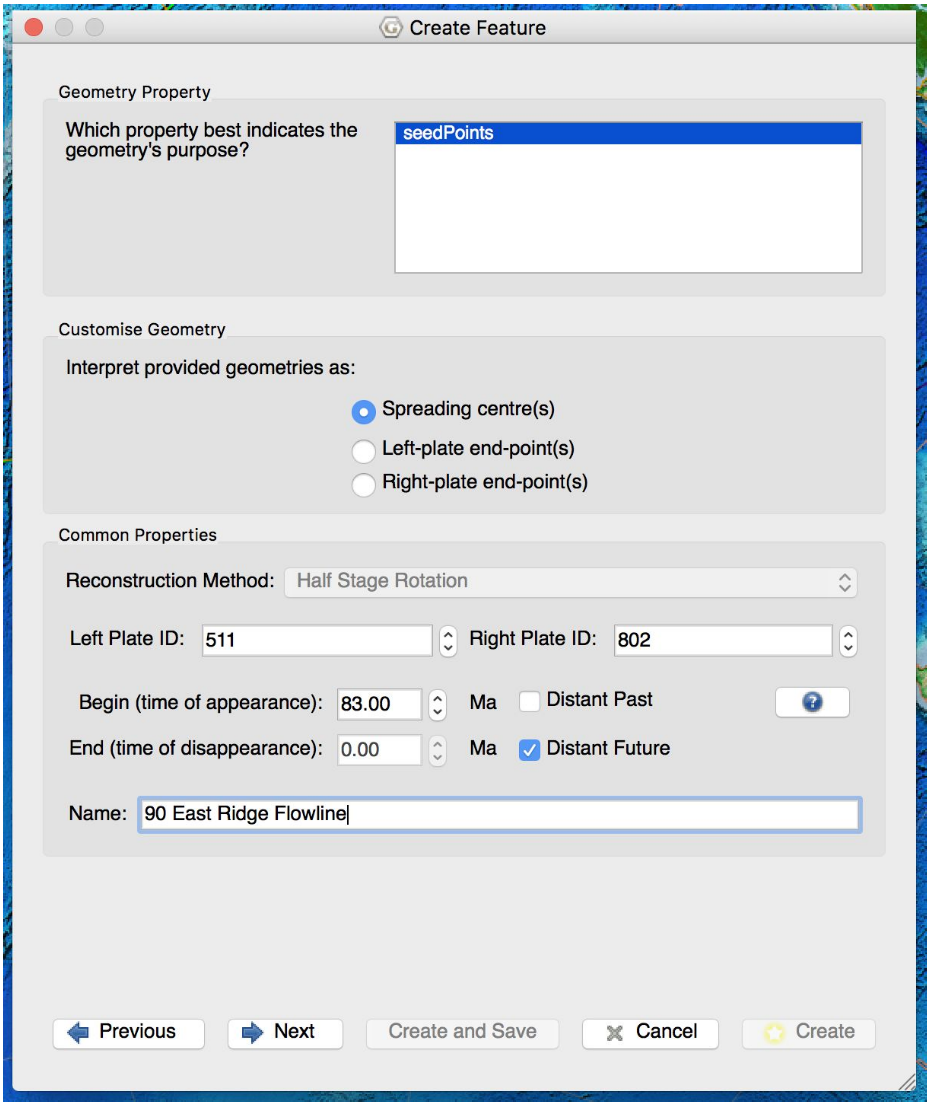


Figure 9. Create Feature menu – flowline properties.

13. In the new menu which appears, under 'Available Properties' select gpml:times and click "+ Add".

14. A new menu will pop down. Under 'Insert multiple times', fill in the following values:

From: 83 Ma
To: 0 Ma
in steps of: 10 Ma

Then click "Insert" (Figure 10).
Click "OK"

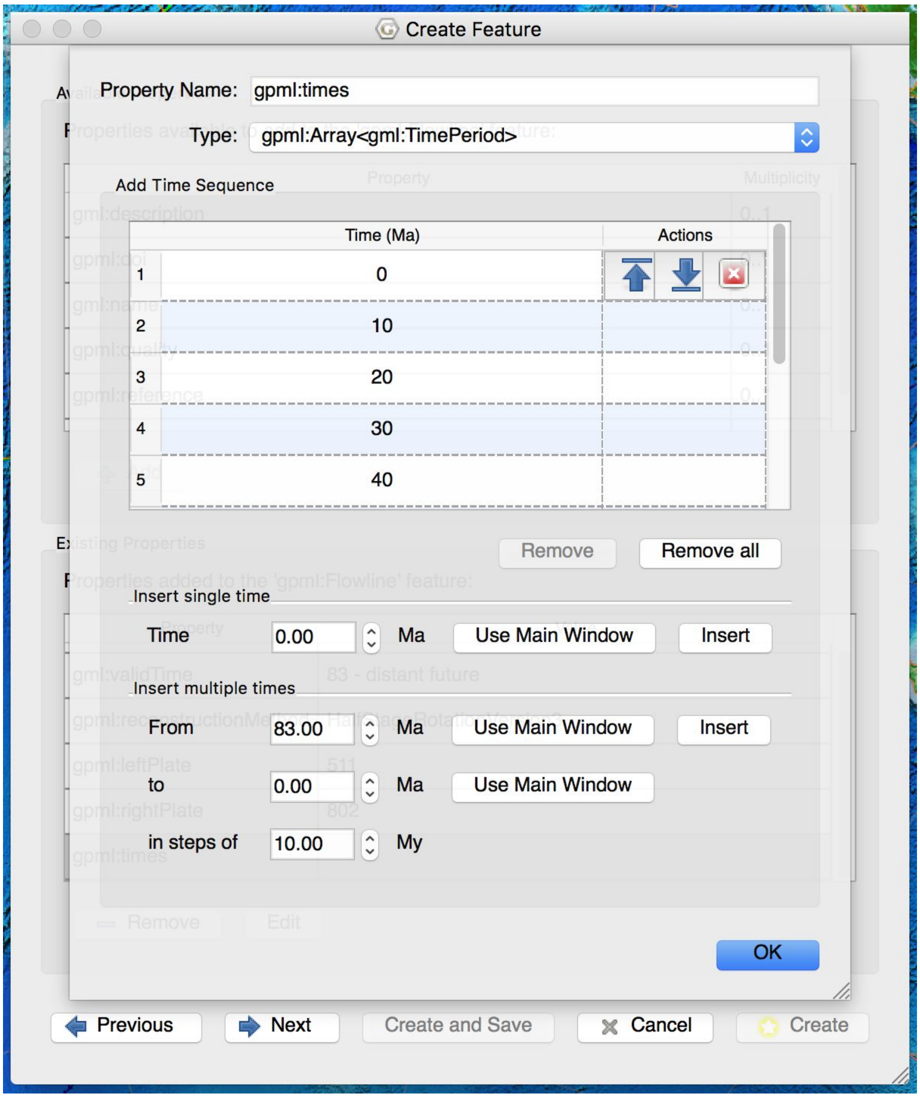


Figure 10. Create Feature menu – inserting time steps.

15. Select "→ Next"

16. Choose '<Create a new feature collection>' then click "Create".

A coloured flowline will have appeared (Figure 11). The arrows indicate the direction of plate motion at the time it appears, with a yellow point indicating the position of the spreading ridge. You should see that the green half of the line flows parallel to the fracture zones as it moves northeast away from the spreading center, and then turns and follows directly along the 90 East Ridge. This alignment suggests that the plume feeding the ridge was interacting with the plate boundaries related to the Indo-Australian-Antarctic triple junction, as it formed (Whittaker et al 2013).

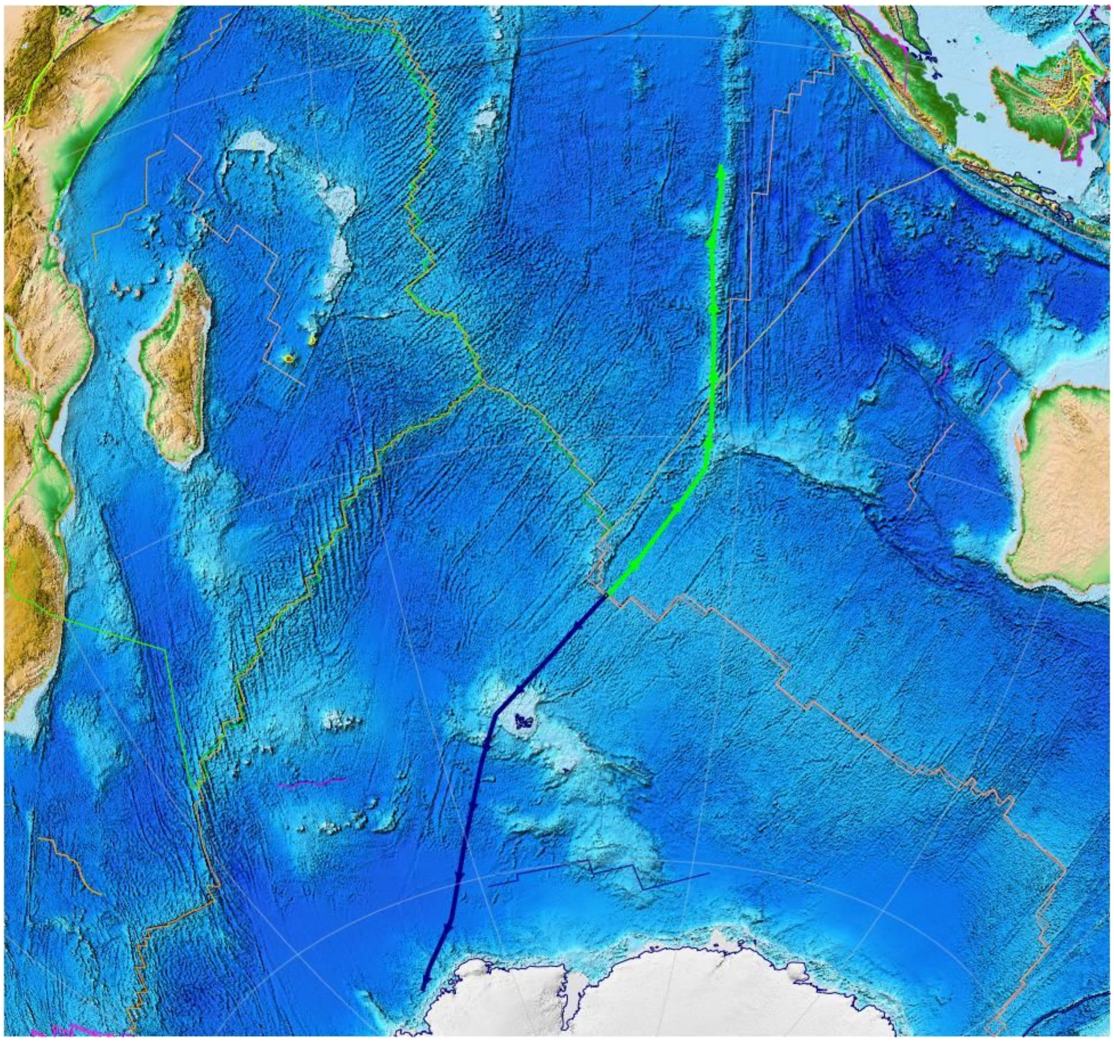


Figure 11. Flowline following the 90 East Ridge.

Play around with reconstructing the flowline back in time. If you use the slider or arrows to move forward in time you will see the flowline as it is created. You may notice that between 83 Ma and 43 Ma the yellow seed point is not constrained to a plate boundary. This is because the spreading ridge on which we set it has not yet formed. If you set the time to 54 Ma and play forward, you will be able to see the formation of the ridge following the fracture zone associated with the India-Australia Mid Ocean Ridge (Figure 12).

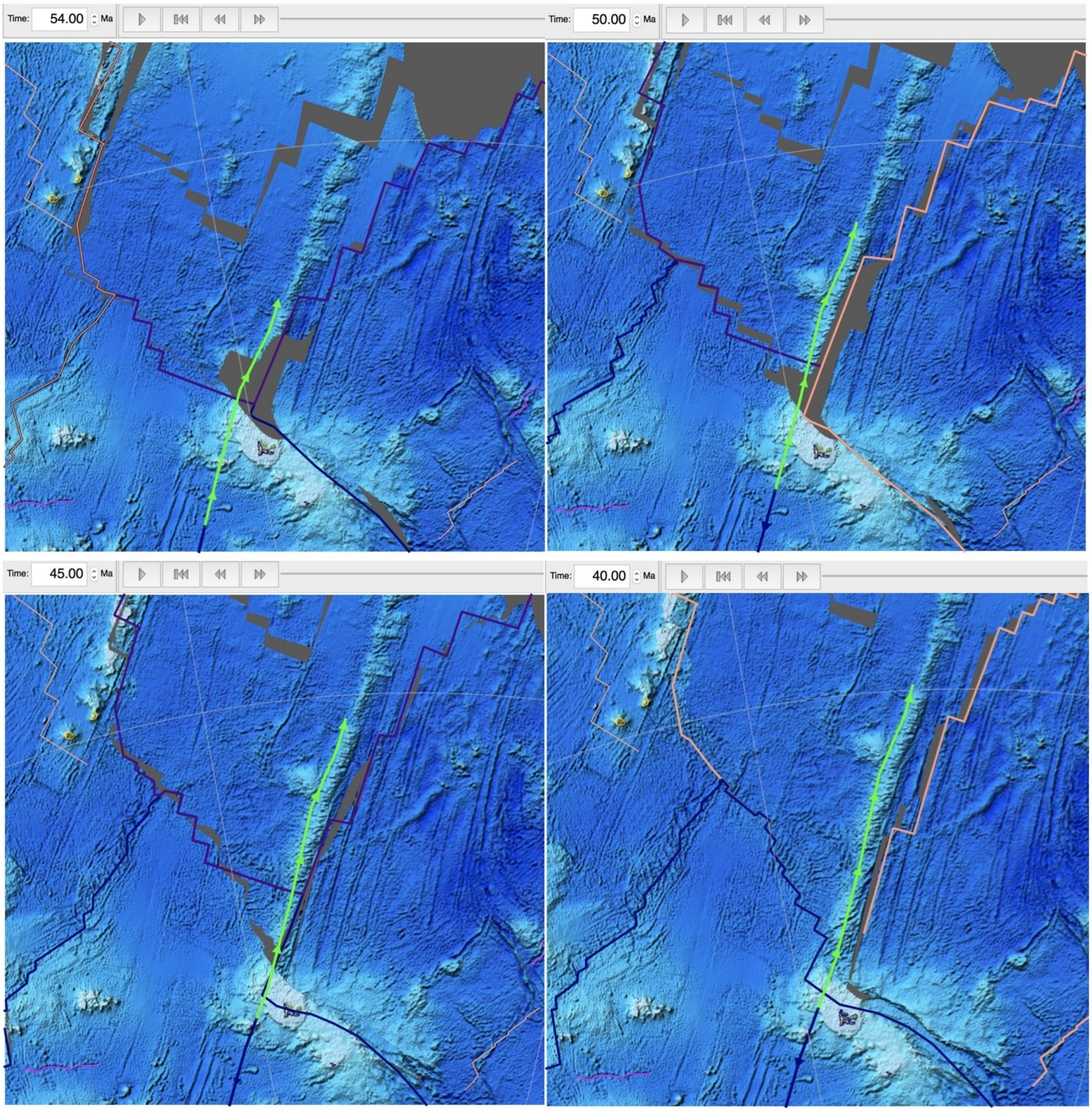


Figure 12. Time progression (54, 50, 45 and 40 Ma) showing the interaction of the 90 East Ridge with a fracture zone. The flowline is in lime green and the fracture zone associated with the Indian-Australian Mid-Ocean Ridge appears either purple or orange.

17. At this point you may like to save your flowline. Click on the  symbol in the bottom right hand corner of the screen. This will bring up the ‘Manage Feature Collections’ window. Scrolling to the bottom of the list of feature collections you will see a layer called ‘New Feature Collection’, highlighted in orange, this is the flowline you just created (Figure 13). For this layer click the ‘Save As’ icon , under ‘Actions’.

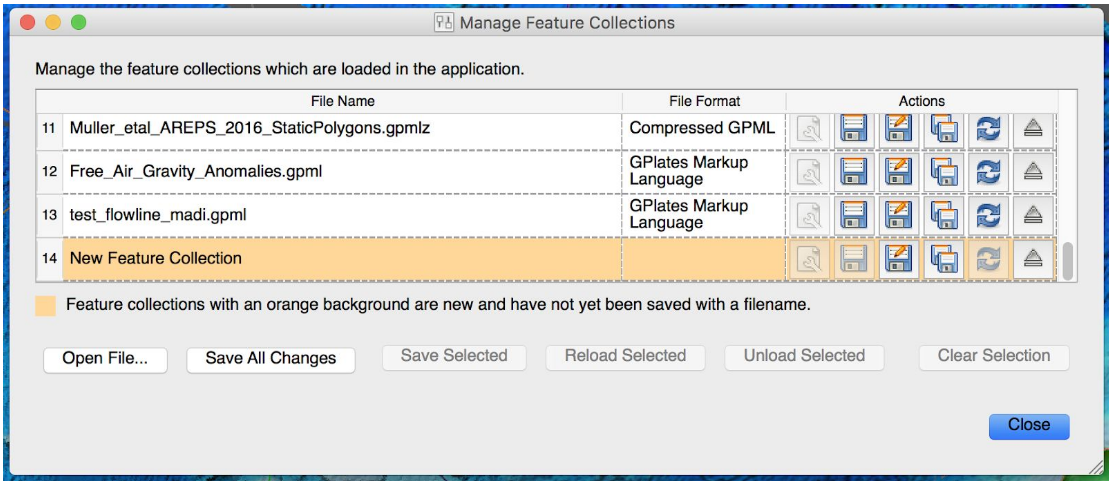


Figure 13. Manage Feature Collections window.

18. Navigate to the tutorial data folder and save your flowline with an appropriate name such as '90 East Ridge Flowline'. Keep the default file extension .gpml. Then click "Save".

19. You may also like to make visible the 'Free Air Gravity' raster we uploaded previously as an additional way to visualize the ridge formation.

Ok, now we are going to compare our flowline, which uses relative motions, with a motion path that will reveal the absolute motion of the volcanic trace relative to the mantle. Motion paths can be used to track the absolute motion of surface features formed by mantle plume hot spots. Comparing the generated path produced by the rotation file to the actual hotspot track can give you a clue as to other tectonic process that may be interacting with the plume, confusing the plume trace and age progression.

20. As was done for the flowline, create a seed point, but this time placing it on the northern region of the Kerguelen Plateau (Figure 14).

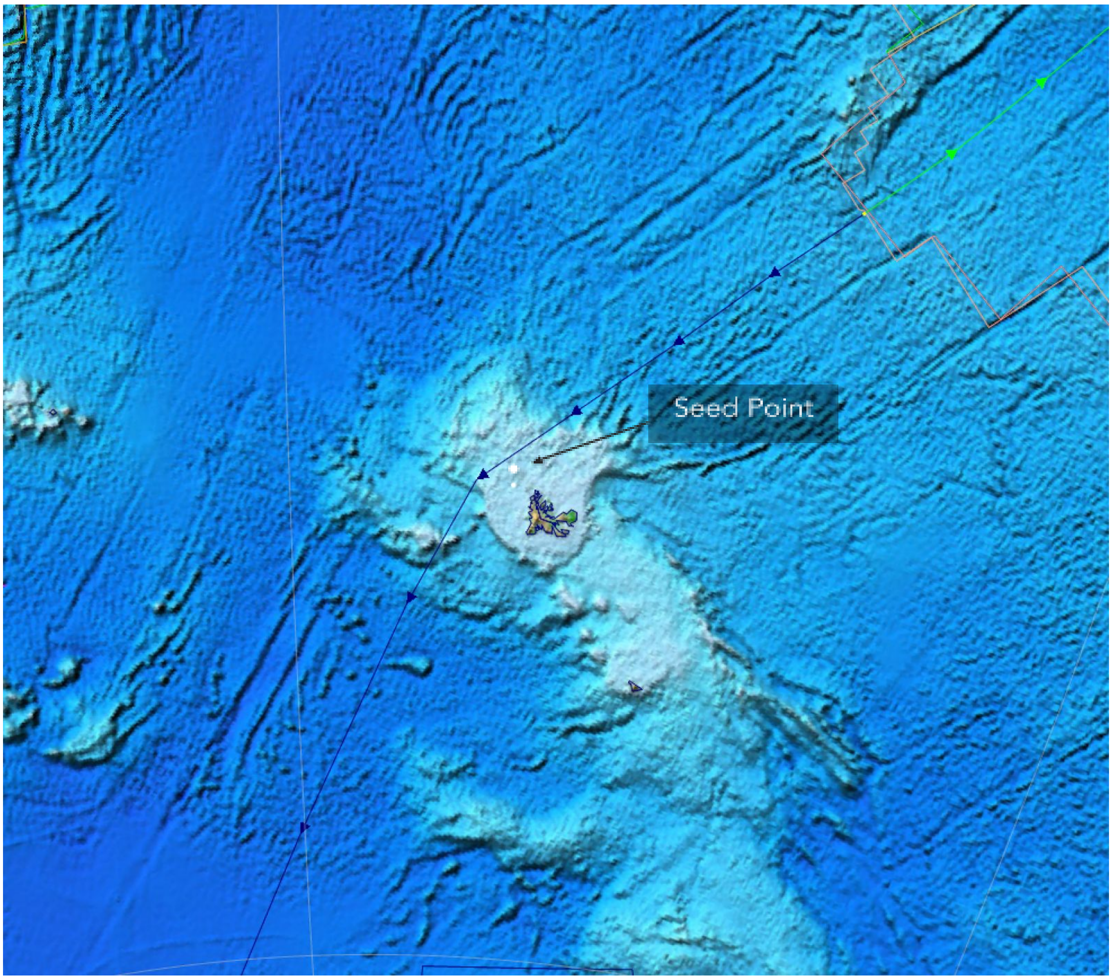


Figure 14. Location of seed point used to create motion path.

21. Once you are happy with its location, click "Create Feature…" in the bottom right hand corner.

22. This time choose your "Feature Type" to be 'MotionPath' and click "Next".

23. Under 'Common Properties', fill in the following fields (Figure 15):

Plate ID: 001 (this is the mantle)
Relative Plate ID: 511 (Central Indian Basin)
Begin (time of appearance): 83 Ma
End (time of disappearance): select 'Distant Future'
Name: 90 East Ridge Motion Path

Click "Next".

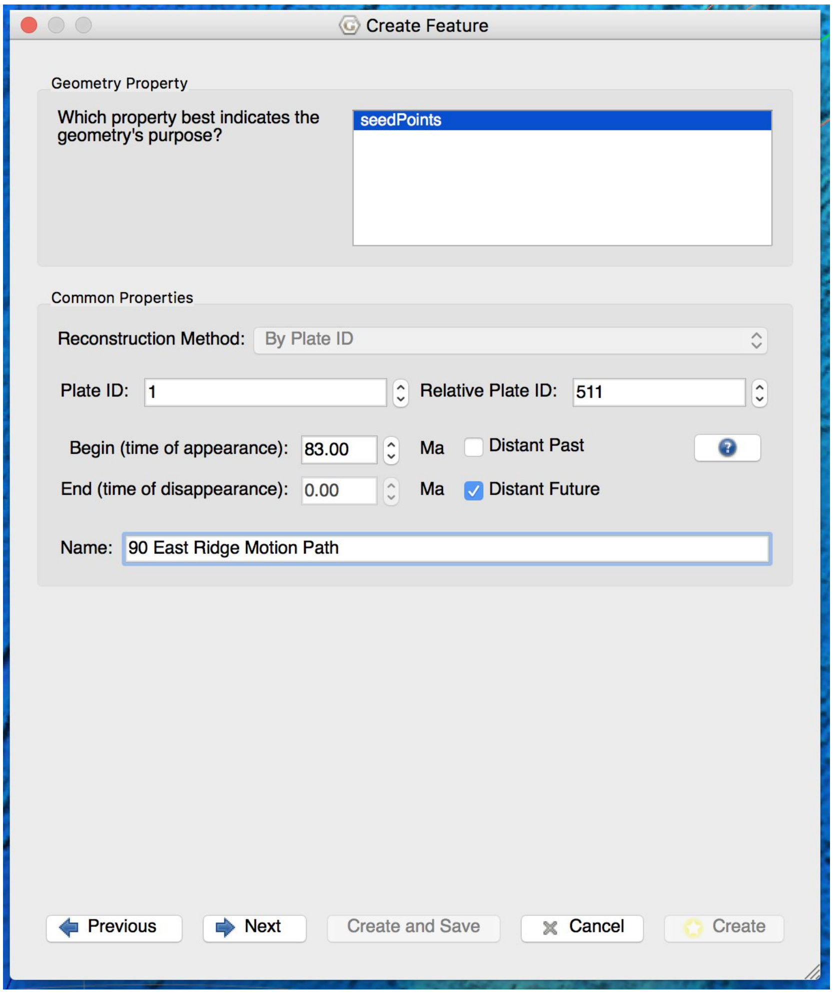


Figure 15. Create Feature menu – motion path properties.

24. The 'gplm:times' property we created for the flowline should automatically be included as part of the 'Existing Properties' list. If not, follow the same process outlined in steps 13 and 14.

25. Click "Next", select '<Create a new feature collection>' then click "Create".

26. A light blue motion path should now have appeared (Figure 16).

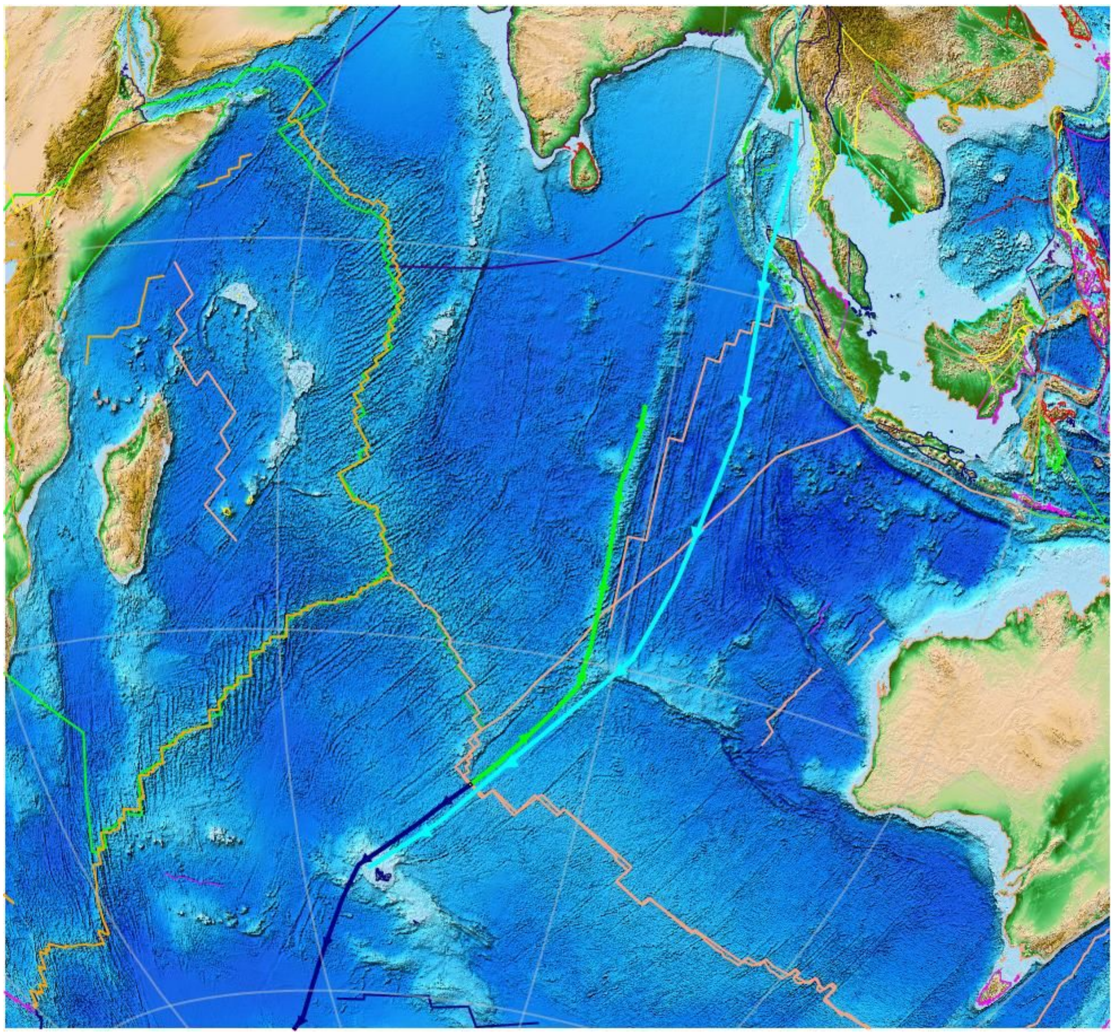


Figure 16. Display of globe showing newly created motion path (light blue) and previously created flowline (black and lime green) at 0 Ma.

27. Take the time to save this feature as we did for the flowline, using an appropriate name, such as ’90 East Ridge Motion Path’.

Time for one last piece of evidence to help us visualize the evolution of the ridge:

28. Locate the ‘Global_Geology.gpml’ file in the tutorial data bundle and drag it onto the globe (Figure 17).

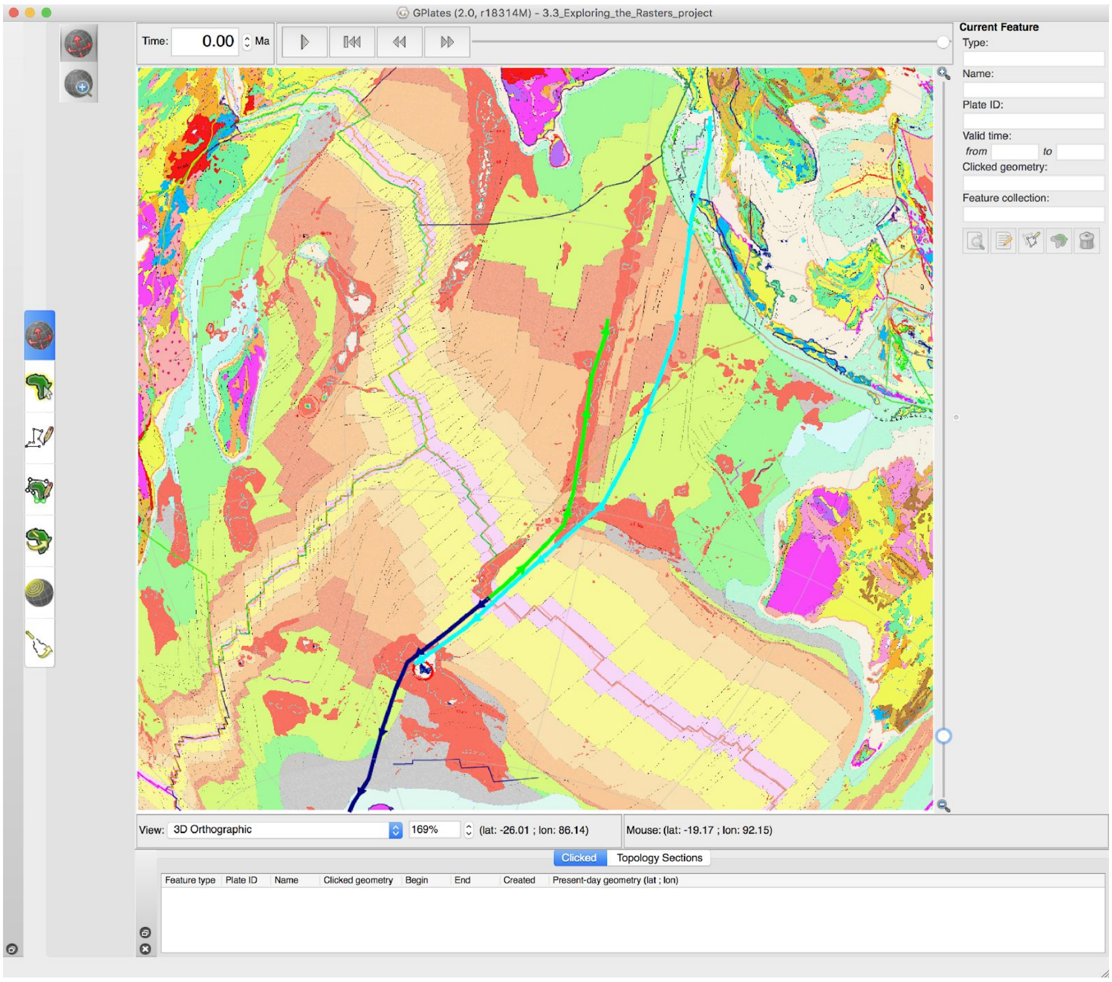


Figure 17. Globe displaying the 'Global_Geolgy' raster at 0 Ma.

This raster displays a range of geological features from major rock facies to glacial extents and large igneous provinces (LIP). Have you noticed that the 90 East Ridge and Kerguelen Plateau (coloured red) are formed from a different rock type than the surrounding sea floor? See if you can locate this rock facies in the legend for this raster (the legend is comprised of two images, included in the data bundle). They are defined as oceanic LIP plateaus, which tells us that they are of volcanic origin.

29. Expand the 'Global_Geology' layer. In the "Inputs" section, click on the "Add new connection" button under "Reconstructed Polygons:" and select the static polygons file from the list. Under "Age grid raster:" again click on the "Add new connection" button and select "Seafloor_Age_Grid". Figure 18 displays how the layer should look after making these changes. This has allowed us to reconstruct this layer back through time.

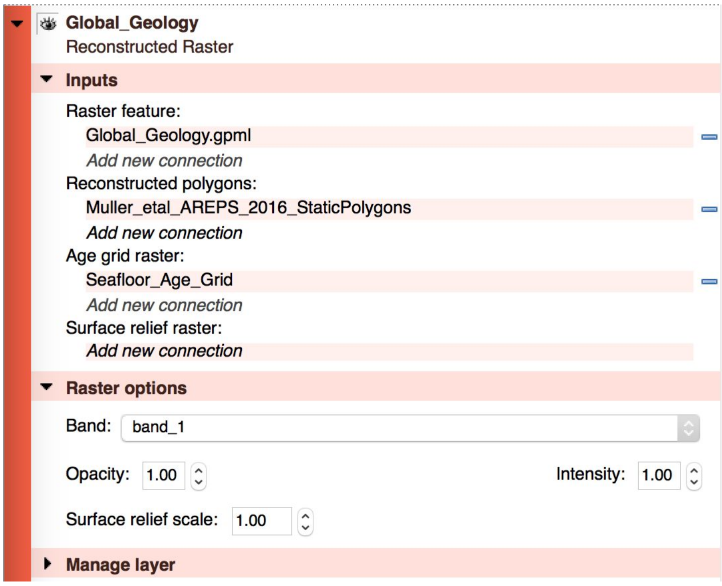


Figure 18. How your 'Global_Geology' layer should look after completing step 29.

Ok, we now have three different rasters loaded to help us visualize the 90 East Ridge and Kerguelen Plateau, and both a flowline and motion path to help track their evolution.
Let's set the time to 83 Ma and reconstruct through to present day. Watch how the development of the flowline and motion path varies, resulting in quite dramatic differences at present day.

Can you see that it is the motion path, which is using absolute motions relative to the mantle, has the worst fit with the ridge? This suggests that the mantle plume has either moved (as a result of mantle wind or the splitting of the plume into a number of diapirs) or that there has been preferential upwelling of the magma along the plate boundaries. Whittaker et al. (2013) found that a fixed hotspot reference frame produced the best match between the reconstructed position of the 90 East Ridge and the inferred track of the plume, suggesting that preferential upwelling is the most plausible explanation of the differences we see in the relative and absolute hot spot paths we created.

Keep GPlates and your generated flowline and motion path open for exercise 2.

Exercise 2 – Exporting Projected Geometries & Rasters

This short exercise follows on directly from Exercise 1 and explains some of the ways that you can export and visualize the rasters and projected geometries.

1. You should still have GPlates open and your files loaded from exercise one. If not, open GPlates and drag and drop the GPlates project file 'Tutorial_3.4.gproj' from the data bundle onto the globe. The files in this project will be sufficient to complete the exercise, but note that figures will look different, as you will have the topography raster displayed.

2. Rotate the globe and zoom in so that you are focused on the 90 East Ridge (Figure 19). Your motion path and the northern segment of the flowline should be visible.

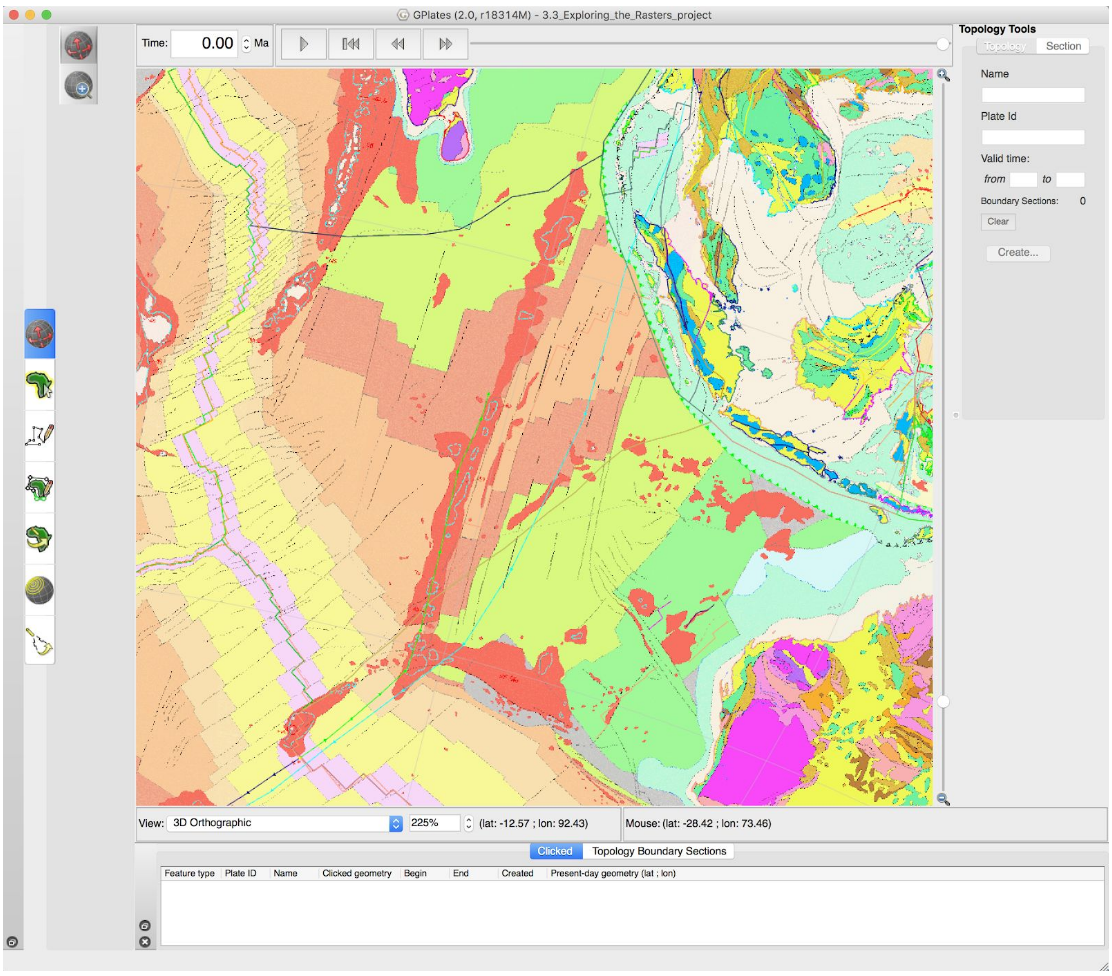


Figure 19. Globe focused on the 90 East Ridge with the global_geology raster visible.

3. Navigate to the 'Export' window using: Reconstruction > Export…(Figure 20).

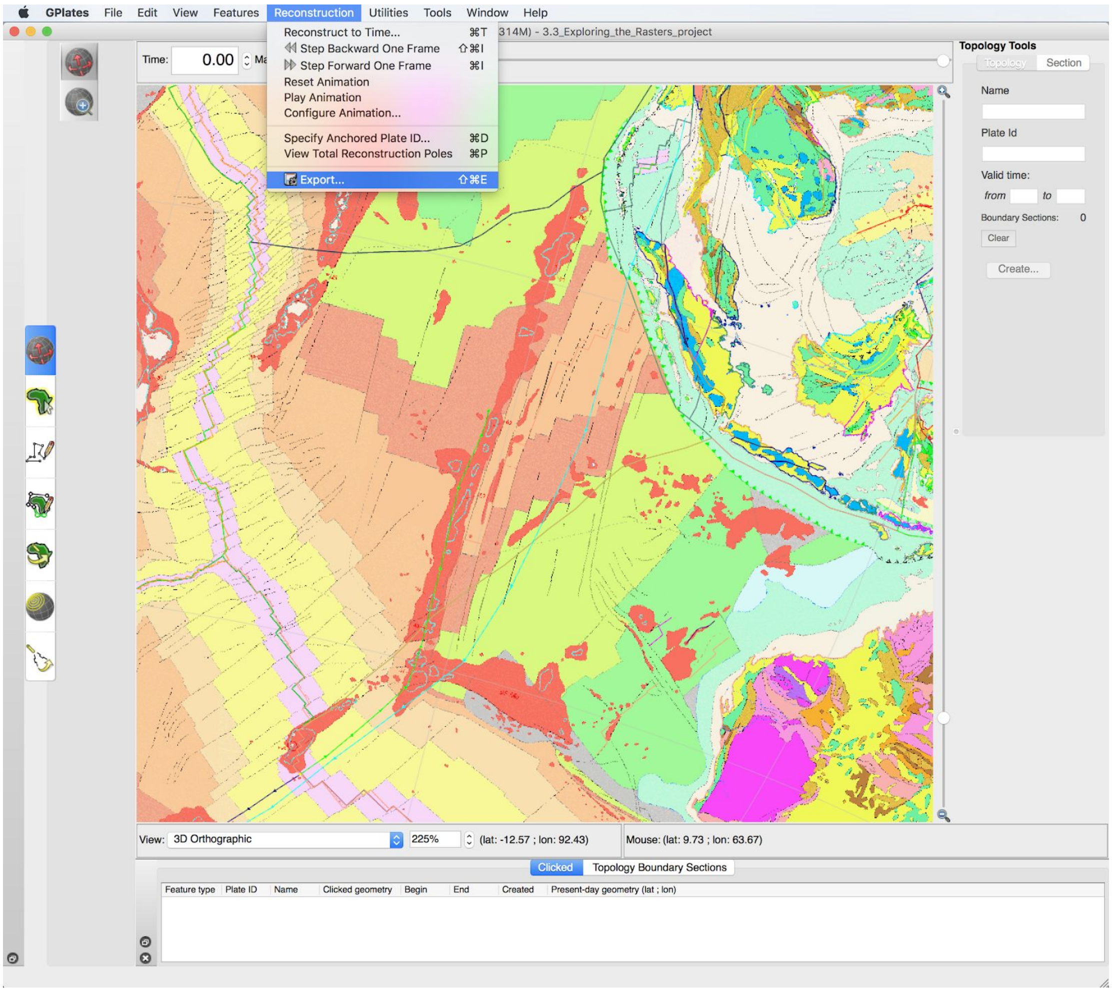


Figure 20. Step 3.

4. At the top of the window that appears, you are given two options:
   - 'Export Time Sequence of Snapshots'
   - 'Export Single Snapshot Instant'

We are going to begin by exporting a single snapshot, so select this option. The window should now appear as it does in Figure 21.

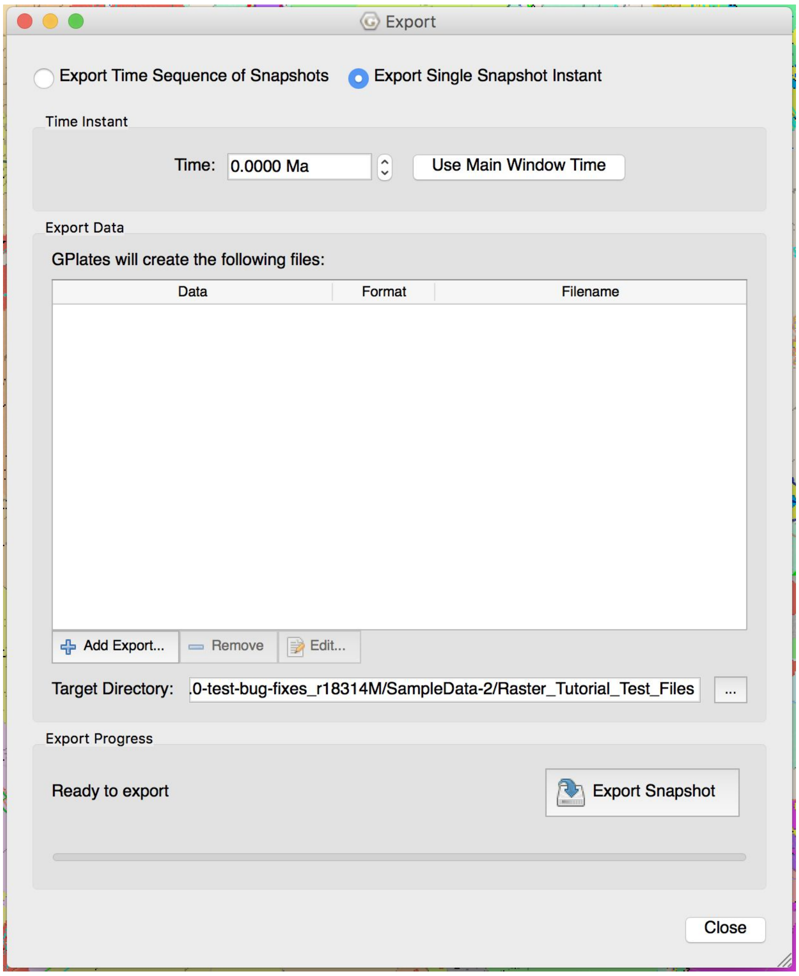
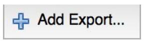


Figure 21. Options for exporting a single snapshot.

5. Leave the default time as 0 Ma.

6. Click the 

button and a new window should appear. Select the flowing for each of the steps and your window should look like Figure 22:

1. Select 'Projected Geometries (and Rasters)'
2. Choose the output file format 'SVG (*.svg)'

3. Under 'Image Resolution', use the main window dimensions by clicking on the  button.

4. Add more detail to the default name given under 'Template:'. For example you might use 'snapshot_90EastRidge%0.2fMa'.

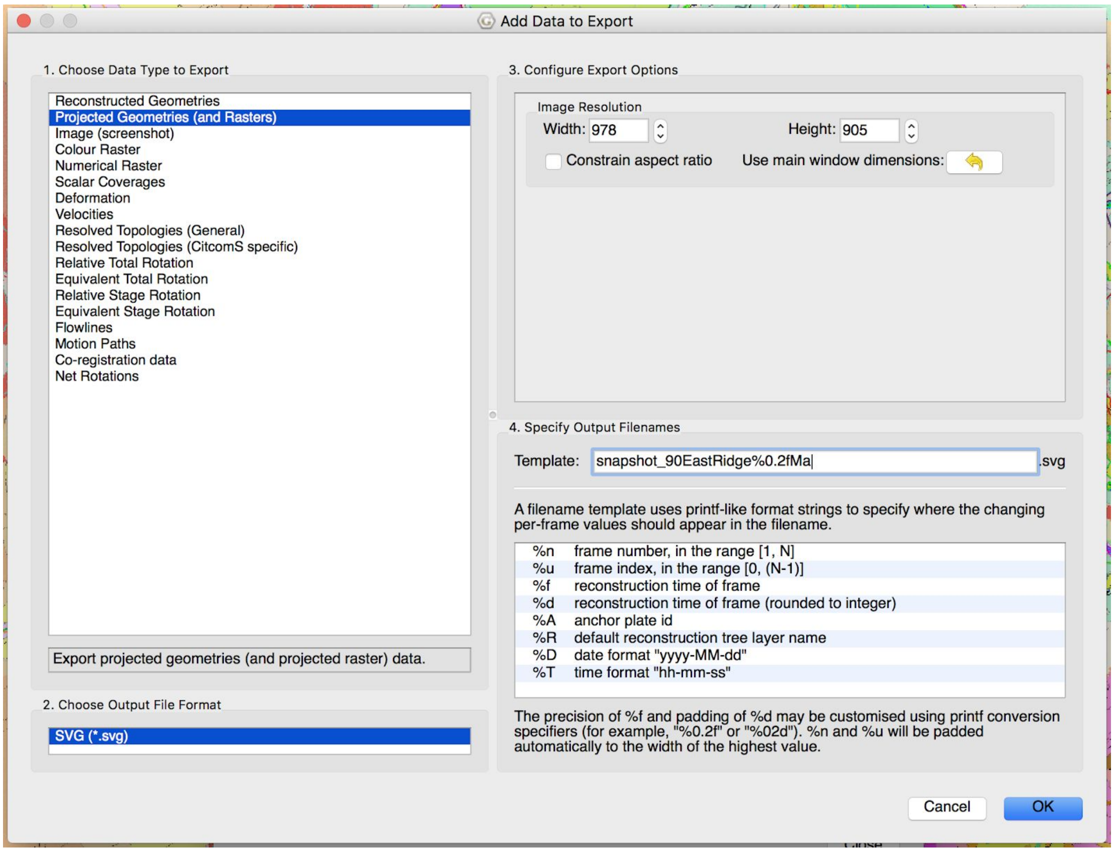
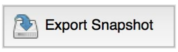


Figure 22. 'Add Data to Export' window with options selected.

7. Click 'OK'

8. A new row of data, highlighted in dark blue should have appeared in the 'Export Data' section. The next step is to specify your target directory by clicking 

 and navigating to the folder you would like to save it in, perhaps the tutorial data bundle folder. Click 'choose'.

9. Now you are ready to export. Click: 

10. The SVG file you have produced will now allow you to manipulate and edit the snapshot. For example, loading the file into Adobe Illustrator will allow you to select individual lines and increase their thickness. A very brief illustration of this is given in Figure 23. Refer to specific tutorials for Adobe Illustrator on more information on how to do this.

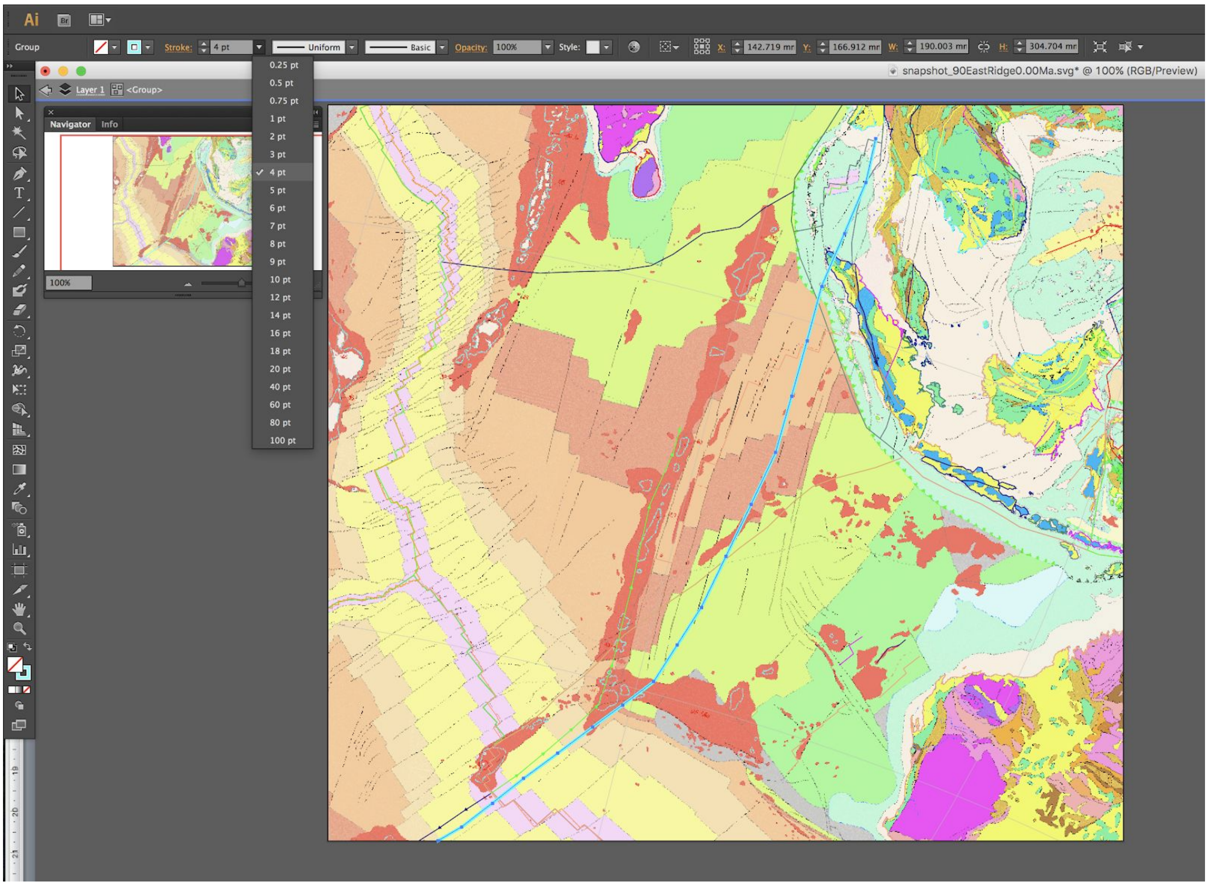


Figure 23. Increasing the thickness of the motion path line using Adobe Illustrator.

11. We are now going to export a time sequence of snapshots. The 'Export' window should still be open, if not, open it using: Reconstruction > Export…

12. Select the 'Export Time Sequence of Snapshots' option (Figure 24).

13. Zoom out so that the entire globe is in view, still focused on the Indian Ocean (Figure 24.)

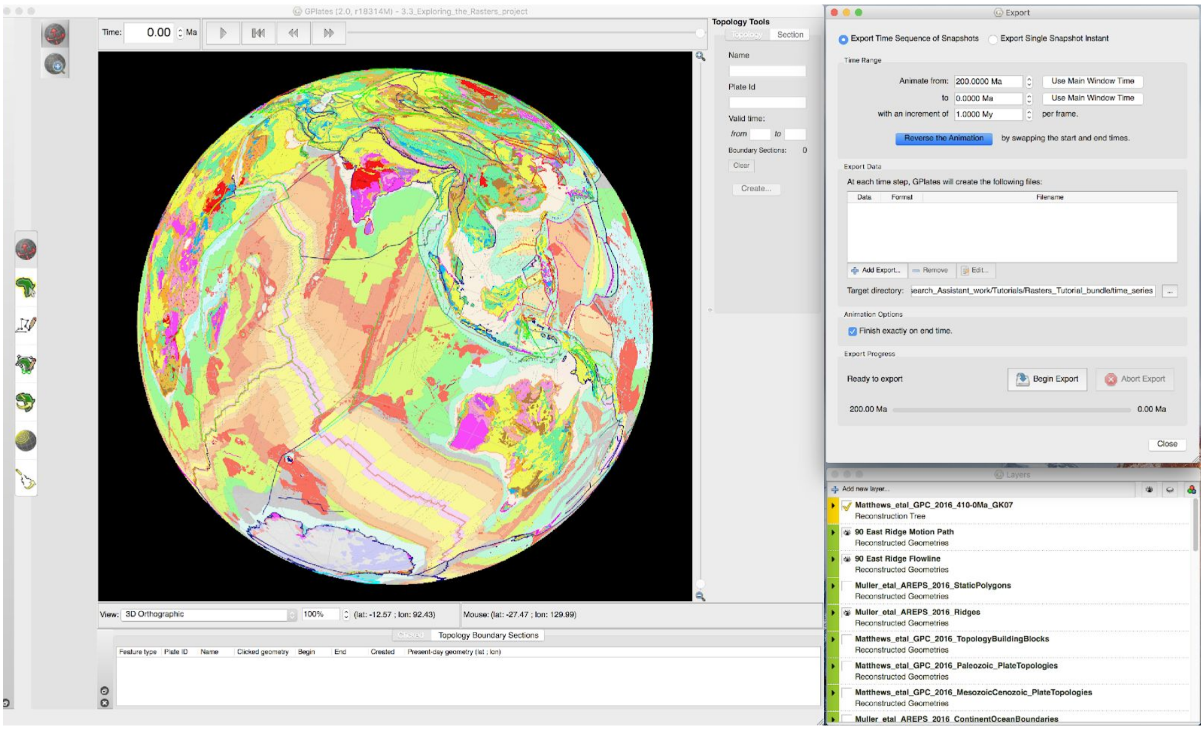
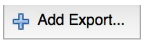


Figure 24. By step 13 you should have your entire globe visible, focused on the Indian Ocean, and the 'Export' window open and ready to input variables.

14. In the 'Export' window under 'Time Range', set the following times:
Animate from: 83 Ma
to: 0 Ma
with an increment of: 1 Ma

15. Click 

, and in new window that appears, select the following for each of the steps (Figure 25):
1. Select 'Image (screenshot)'
2. There are a number of good options, but in this tutorial we will select 'Portable Network Graphics (*.png)' as the file format.
3. Under 'Image Resolution', use the main window dimensions by clicking on the  button.
4. Add more detail to the default name given under 'Template:'. For example you might use 'Indian_Ocean_Reconstruction_%0.2fMa'. The '%0.2f' ensures that each file is labeled with the correct reconstruction time.

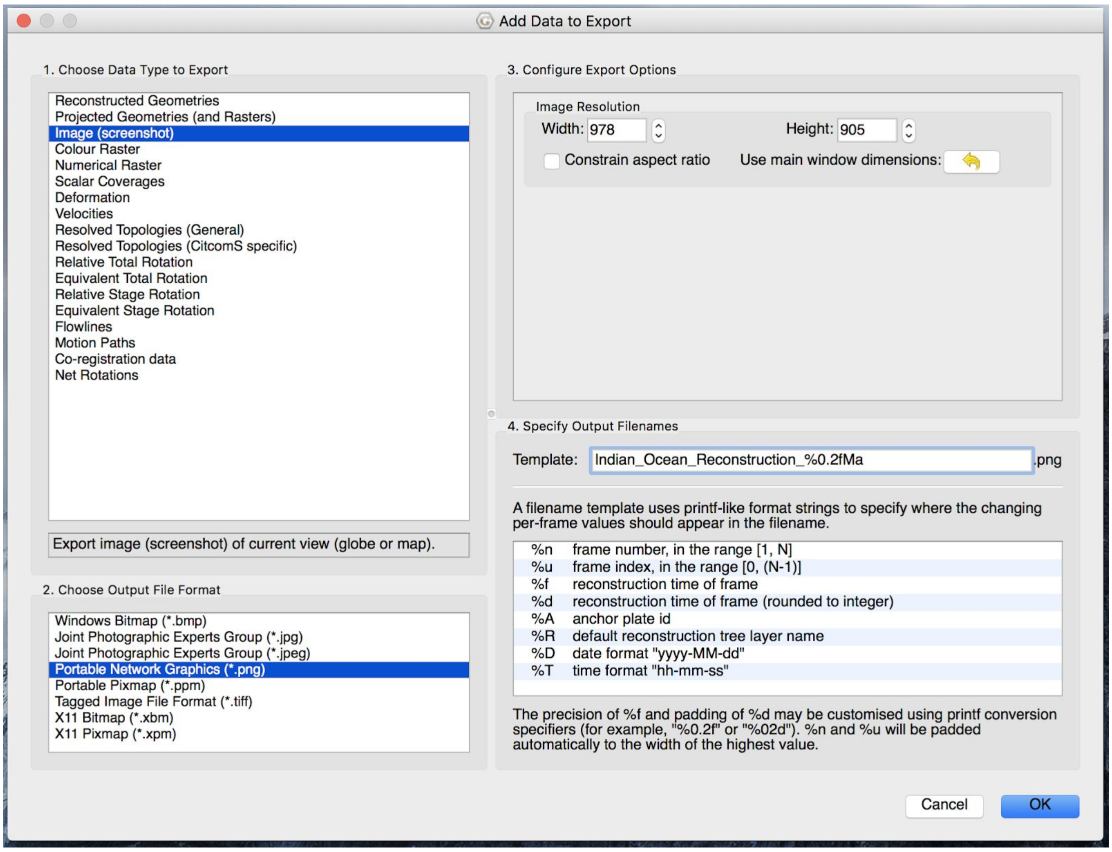


Figure 25. Step 15.

16. Click ‘OK’

17. A new row of data, highlighted in dark blue should have appeared in the ‘Export Data’ section. The next step is to specify your target directory by clicking  and navigating to the folder you would like to save it in. It is best to create a new folder for this step, as 83 files will be produced. Once you have navigated to the desired folder click ‘choose’. Your export window should look as it does in Figure 26.

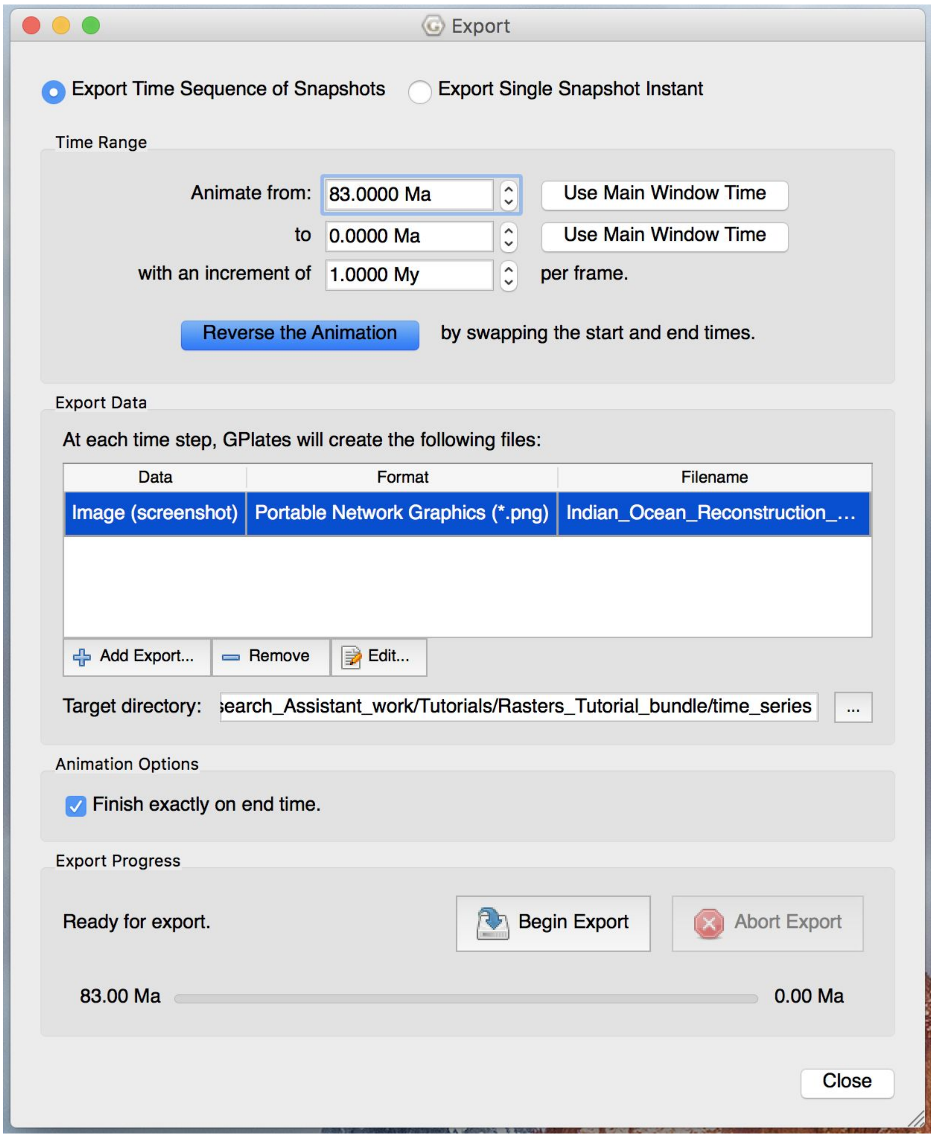


Figure 26. Step 17.

18. You are now ready to export the snapshots. Click: 

. Generating the files will take a few seconds.

19. In the selected folder you should now see a .png file for each time step (Figure 27 gives an example for 24 Ma). These images can be used to generate figures or combined to create an animation showing the evolution of the Indian Ocean through time. This tutorial will not go through this step, but some simple programs that could be used to create an animation include: iMovie, Windows Movie Maker, Time Lapse Assembler…

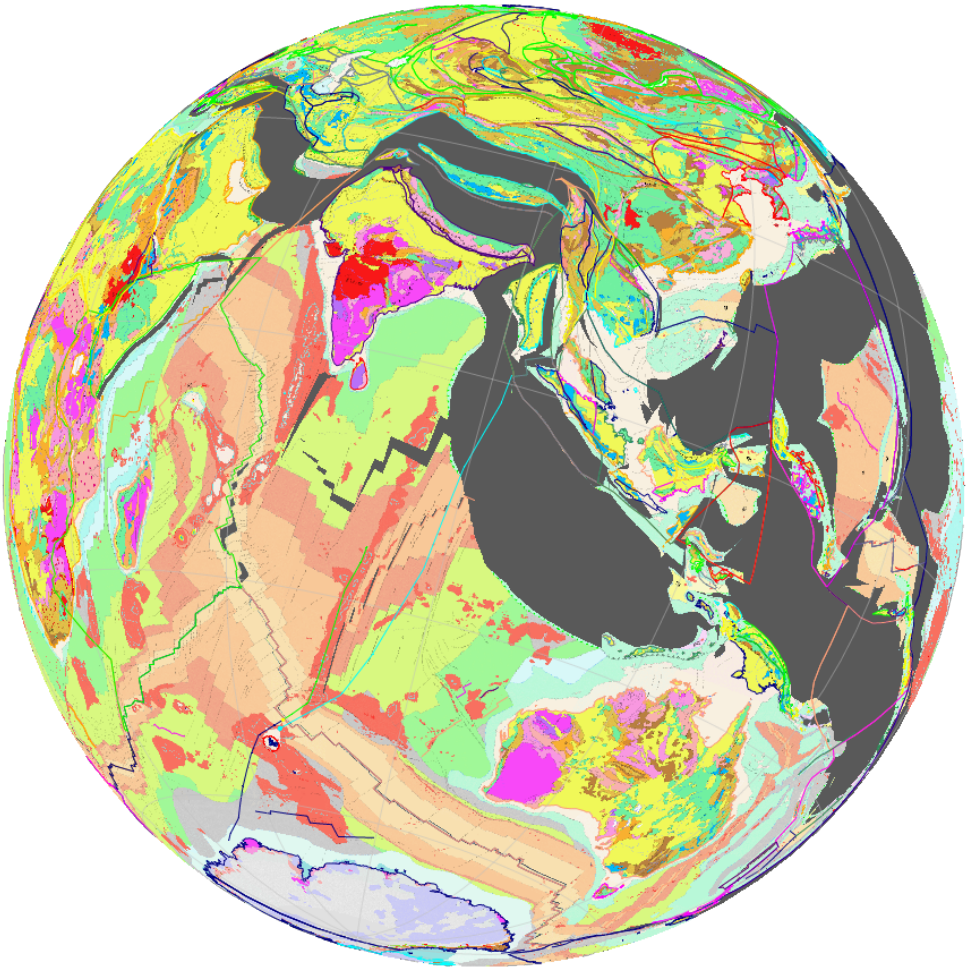


Figure 27. png file showing the arrangement of continents around the Indian Ocean at 20 Ma, visualized using the global_geology raster. The motion path and flowline constructed in the first exercise are also visible.

References

Gibbons, A.D., Whittaker, J.M. and Müller, R.D., 2013. The breakup of East Gondwana: assimilating constraints from Cretaceous ocean basins around India into a best-fit tectonic model. *Journal of geophysical research: solid earth*, *118*(3), pp.808-822.

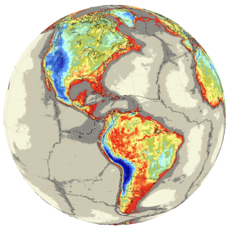
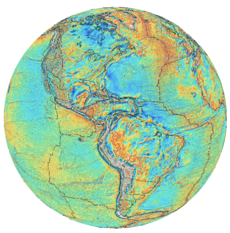
Whittaker, J.M., Williams, S.E. and Müller, R.D., 2013. Revised tectonic evolution of the Eastern Indian Ocean. *Geochemistry, Geophysics, Geosystems*, *14*(6), pp.1891-1909.

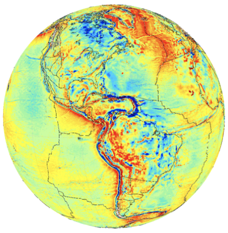
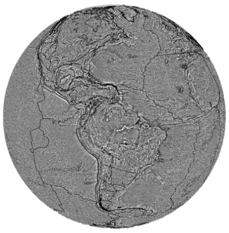
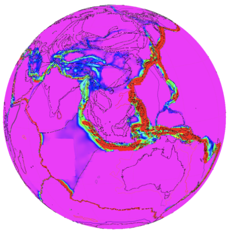
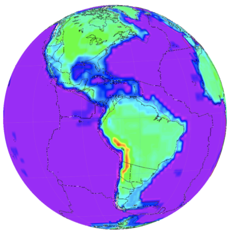
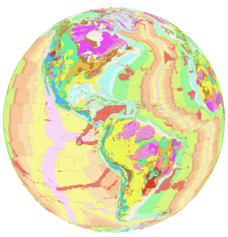
O'Neill, C., Müller, D. and Steinberger, B., 2005. On the uncertainties in hot spot reconstructions and the significance of moving hot spot reference frames. *Geochemistry, Geophysics, Geosystems*, *6*(4).

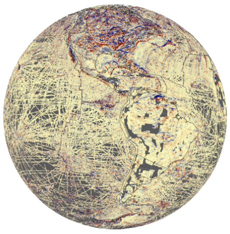
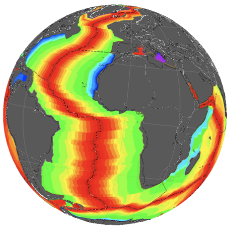
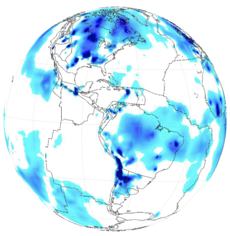
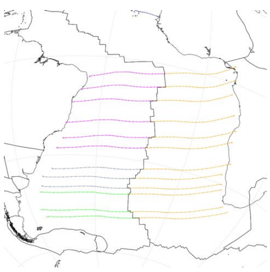
Appendix

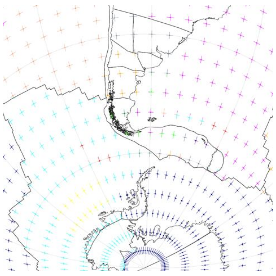
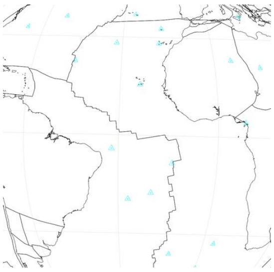
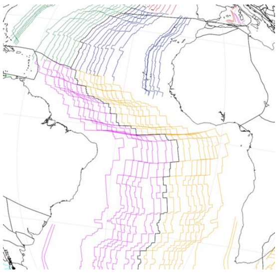
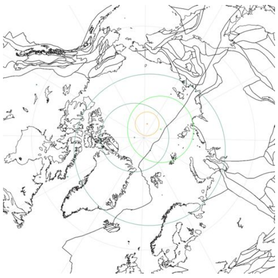
The table below lists the rasters and sample data available in the GPlates 'SampleData' bundle. This can be found here on the GPlates website, or directly downloaded as a zip file here. It is also included in the GPlates 2.1 package application download.

References for each raster can be found in the '_README.txt' file in the Rasters folder.

Raster/SampleData	Path & File Name	File Type	Explanation	Units
Bouguer Gravity Anomalies 	SampleData > Rasters > Bouguer_Gravity_Anomalies > Bouguer_Gravity_Anomalies_WGM.gpml	Colour grid	All gravity anomalies highlight variations in the strength of Earth's gravitation force at a particular location and provide clues about the structure of the lithosphere and mantle. This Bouguer Gravity Anomaly has been corrected for the elevation of the measurement site (free-air correction) and the influence of topography and terrain (Bouguer and terrain corrections). The Bouguer correction accounts for the gravitational attraction of the rocks between the reference ellipsoid (usually sea level) and the measurement elevation. The terrain correction takes into account the fact that Earth's surface is not an infinite horizontal plane (Fowler, 1990). This raster is useful in understanding subsurface density variations (Balmino et al., 2012).	mGal (milligal)
Free Air Gravity Anomalies 	SampleData > Rasters > Free_Air_Gravity_Anomalies > Free_Air_Gravity_Anomalies.gpml	Colour grid	This raster represents gravity anomalies at Earth's surface, accounting for the elevation at which the measurement was taken. Known as the 'free-air' correction, this compensates for the decrease in gravitational force with distance from the Earth's surface. This calculation assumes that any material between the measurement elevation and sea level is simply 'free-air' (Fowler, 1990).	mGal (milligal)

Isostatic Gravity Anomalies 	SampleData > Rasters > Isostatic_Gravity_Anomalies > Isostatic_Gravity_Anomalies.gpml	Colour grid	This Isostatic Gravity Anomaly grid has taken into account isostatic corrections according to the airy isostatic compensation model using a uniform compensation depth of 30 km (Balmino et al., 2012).	mGal (milligal)
Vertical Gravity Gradient 	SampleData > Rasters > Vertical_Gravity_Gradient > Vertical_Gravity_Gradient.gpml	Colour grid	A vertical gravity gradient from Sandwell and Smith (2014). This is a satellite-derived gravity model. It provides extremely useful in locating tectonic structures, and mapping regions where topography remains unmapped, such as deep ocean basins or areas covered by thick sediments.	Gravity VGG (eotvos)
Crustal Strain 	SampleData > Rasters > Crustal_Strain > Crustal_Strain.gpml	Colour grid	This raster presents the second invariant of the model strain rate field from Kreemer et al (2003). As evident in the raster, crustal strain is concentrated around major plate boundaries, and provides particular insight into deforming plate boundary zones, and earthquake potential.	Second Invariant Strain Rate [1 x 10^{-9} yr^{-1}]
Crustal Thickness 	SampleData > Rasters > Crustal_Thickness > Crustal_Thickness.gpml	Colour grid	A raster depicting the crustal thickness model (CRUST 2.0) from Laske et al. (2000). This is the thickness of the uppermost, solid layer of the Earth, which extends down to the crust mantle boundary (also known as the *Mohorovičić discontinuity*, or *Moho).*	km
Global Geology 	SampleData > Rasters > Global_Geology > Global_Geology.gpml	Colour grid	This raster displays a world geological map from Bouysse (2014). This map provides information on major rock facies and the relative geological Eras in which they formed. It also indicates the locations of major plate boundaries, large igneous provinces, glacial extents, and a range of other geological features.	Colour and pattern labeling scheme

Magnetic Anomalies (EMAG2) 	SampleData > Rasters > Magnetic_Anomalies_EMAG2 > Magnetic_Anomalies_EMAG2.gpml	Colour grid	A colour grid of magnetic anomalies from EMAG2 (Maus et al., 2009). A good representation of raw magnetic data as directional gridding is not used to fill the gaps. Distinct patterns of magnetic anomalies can inform us of a number of tectonic processes including the formation and temporal evolution of oceanic crust.	nT (nanotesla)
Seafloor Age Grid 	SampleData > Rasters > Seafloor_Age_Grid > Seafloor_Age_Grid_Project.gproj	NetCDF numerical grid	A numerical seafloor age grid with 6 arc minute resolution. This raster clearly depicts the process of seafloor spreading and production of oceanic crust at mid-ocean ridge spreading centers. It also highlights the location of the oldest oceanic crust, within the Mediterranean.	Ma (millions of years)
Slab Age derived from the MIT-P P-wave seismic tomography model. 	SampleData > Rasters > Time-dependent raster sequences > MIT-P08-Asia-UM30-LM12.gpml	Time-dependent raster sequence	A time-dependent raster of slab age, coded from the MIT-P P-wave seismic tomography (Li et al., 2008), where slabs are assumed (on the first order) to sink vertically with a constant sinking rate. The sinking rate applied here is 3 cm/yr in the upper mantle, and 1.2 cm/yr in the lower mantle. This raster can provide insight into the locations of major paleo-subduction zones.	% (Seismic Velocity Anomaly)
Topography 	SampleData > Rasters > Topography > Topography.gpml	Colour grid	A global relief model of present-day land topography and ocean bathymetry, in 1-arc minute resolution. The white regions represent ice sheets. The locations of high mountain ranges and deep-sea trenches are easily explored in this raster.	m (meters of elevation or depth)
Flowlines 	SampleData > FeatureCollections > Flowlines > AtlanticFlowlines.gpml	GPlates geometry file	Flowlines reflect the relative motion between two plate pairs along a mid ocean ridge. For more information on flowlines, refer to tutorial under ‘Reconstructions’ titled ‘Tutorial 2.3: Flowlines and Motion Paths’.	N.A.

Gridmarks 	SampleData > FeatureCollections > GridMarks > Matthews_etal_GPC_2016_GridMarks.gpmlz	Compressed GPlates geometry file	A global set of gridmarks.	5 degrees
Hotspots 	SampleData > FeatureCollections > Hotspots > Hotspots_Compilation_Whittaker_etal_SYMBOLS.dat	PLATES4 GPlates compatible geometry file	Compilation of present day surface hotspot/plume locations from Whittaker et al. (2013). Represented by points (triangular symbols) and split into Pacific and Indo/Atlantic domains.	N.A.
Isochrons 	SampleData > FeatureCollections > Isochrons > Muller_etal_AREPS_2016_Isochrons.gpmlz	Compressed GPlates geometry file	Ocean Floor Isochron Dataset from Müller et al. (2016). Each isochron line joins points on the seafloor of equal age.	Gee and Kent (2007) timescale
Palaeomagnetism 	SampleData > FeatureCollections > Palaeomagnetism > gpml > - Europe2004_RM_Npoles.vgp.gpml - Gondwana2010_RM_NPoles.vgp.gpml - Laurussia2010_RM_NPoles.vgp.gpml - NorthAmerica2004_RM_Npoles.vgp.gpml	VGP GPlates geometry file	Four sample palaeomagnetic data sets from the IAGA Global Paleomagnetic Database. For more information on these data sets, refer to the tutorial under ‘Reconstructions’ titled ‘Tutorial 2.5: Working with Paleomagnetic Data’.	Carboniferous to present